Using Constrained Hydro Simulations to Explore the Gas Component of the Cosmic Web

Renjie Li 2023.06

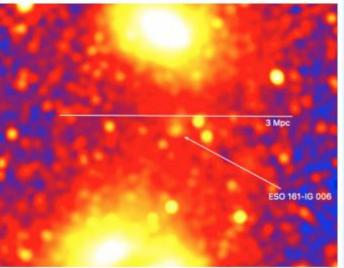
In Collaboration with: Huiyuan Wang, Houjun Mo, Weiguang Cui, Xiong Luo



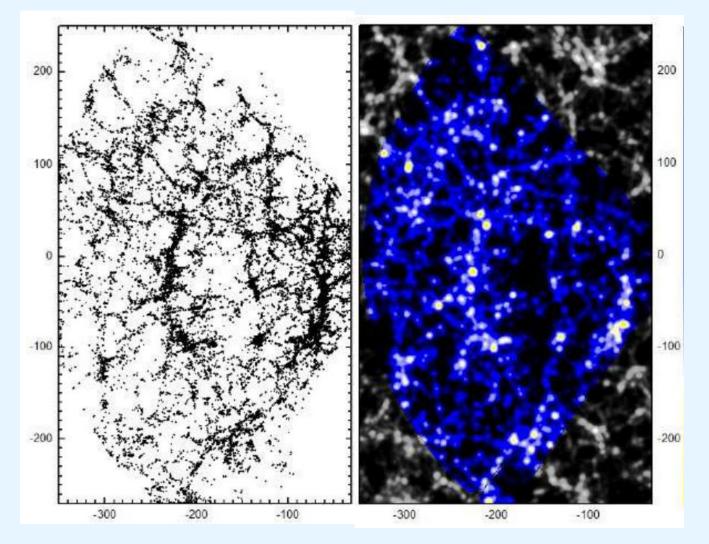
Missing baryon and observation

- Missing baryon problem
- Observation
 - X-ray, SZ effect, absorption line, etc.
 - Gas in filaments
 - Massive filaments between pairs of massive merging clusters (e.g. Planck Collaboration et al. 2013b; Sugawara et al. 2017; Reiprich et al. 2021)
 - Stacking a large number of galaxy pairs (de Graaff et al. 2019)

Reiprich 2021, eROSITA

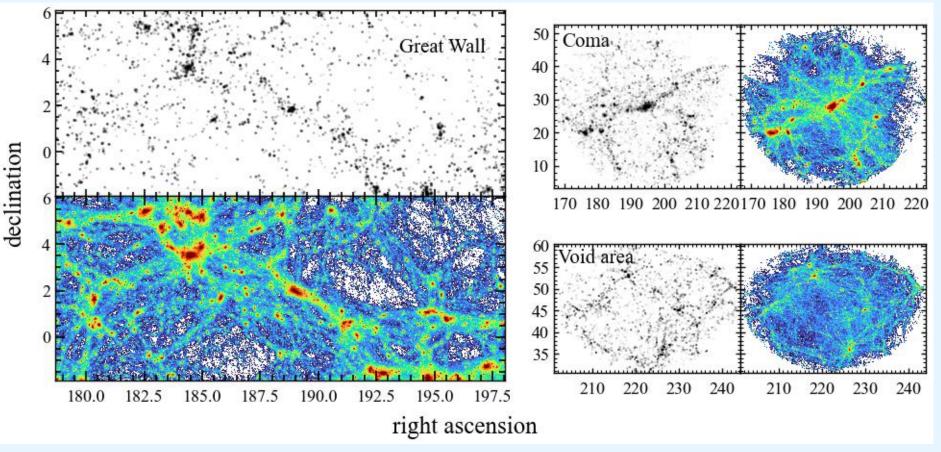






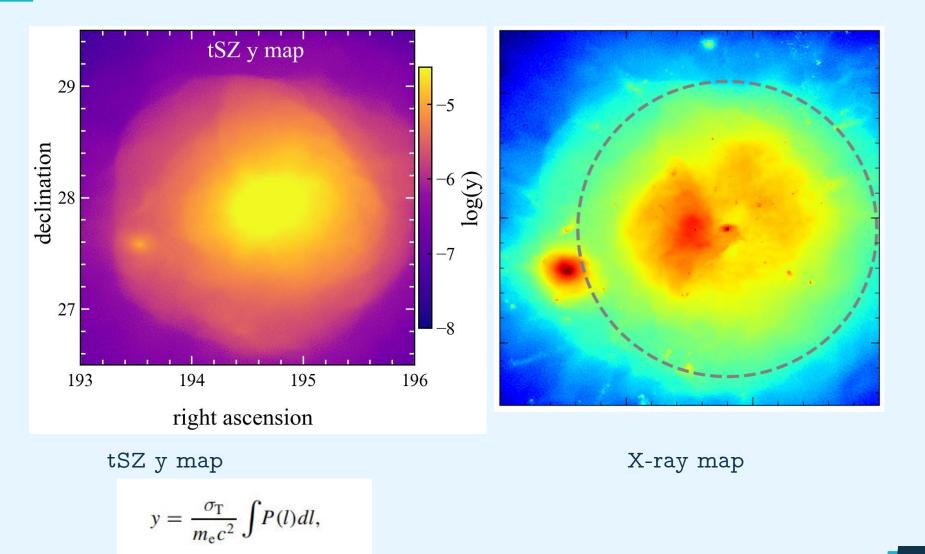
Wang et al. 2016, ELUCID



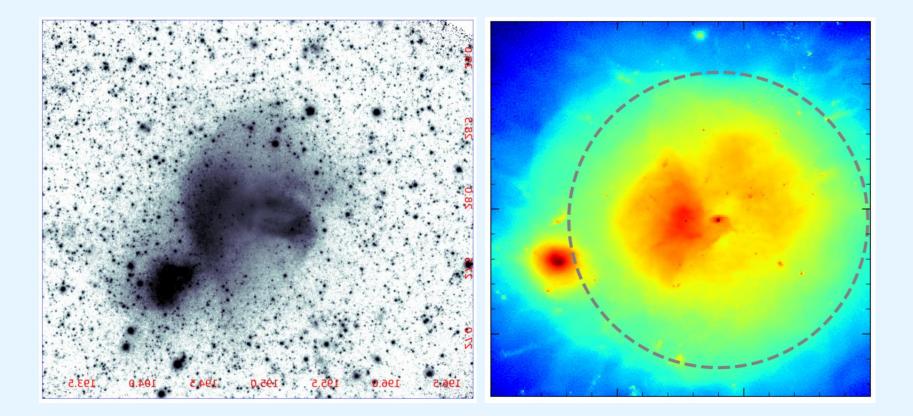


- Gadget-3 code as described in Huang et al. (2019, 2020).
- Feedback from active galactic nuclei is NOT included in our simulations.

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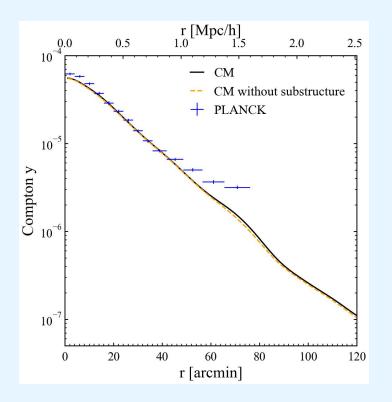
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Churazov et al. 2021, eROSITA



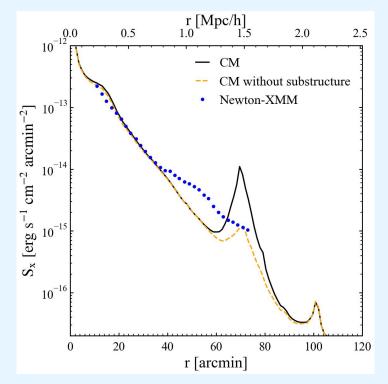
• The tSZ effect



 The data points, taken from (Mirakhor & Walker 2020, private communication)

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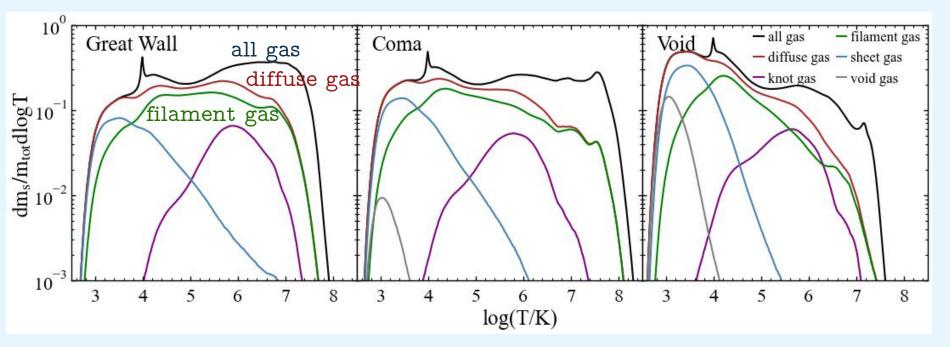
• X-ray



 The data points are based on Newton-XMM observation (Mirakhor & Walker 2020, private communication)

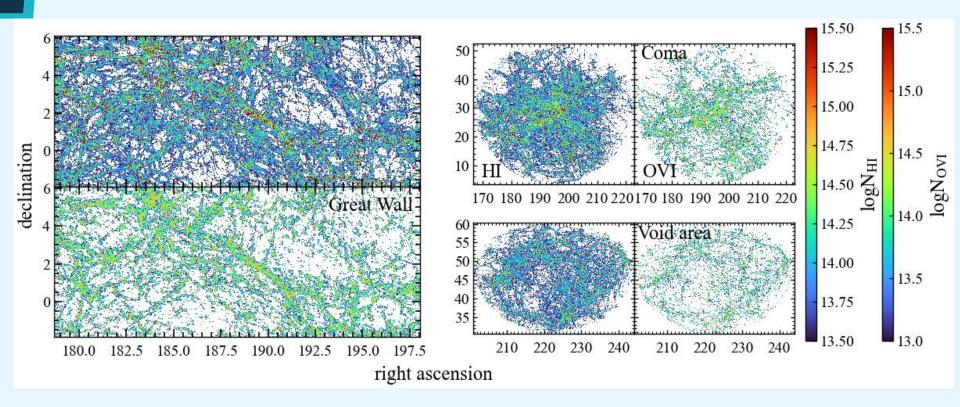
Warm-hot gas in filaments

• The temperature distribution



• Most of the warm-hot diffuse gas resides in filaments.

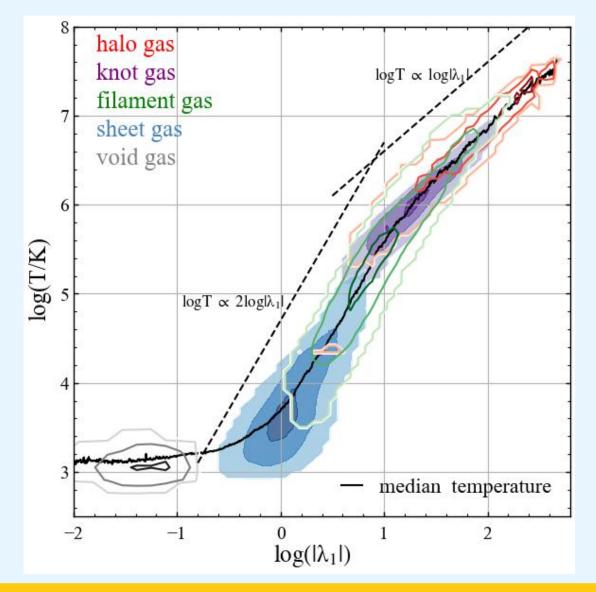
Detect gas in filaments



Low N_{HI} Ly α absorption and most of the OVI absorption with log $N_{OVI} > 13$ are good tracers of the diffuse gas in filaments. (Bradley et al. 2022)



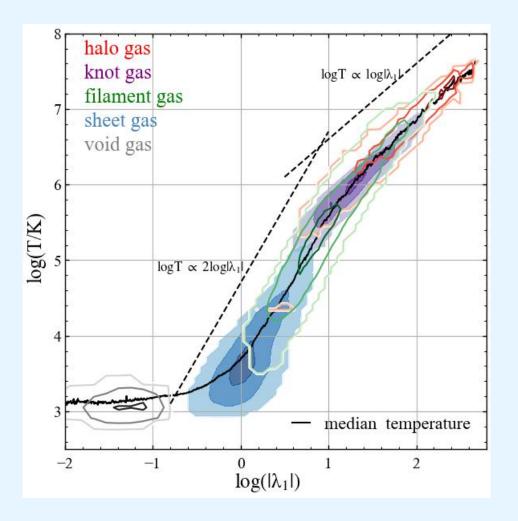
$T-\lambda$ correlation

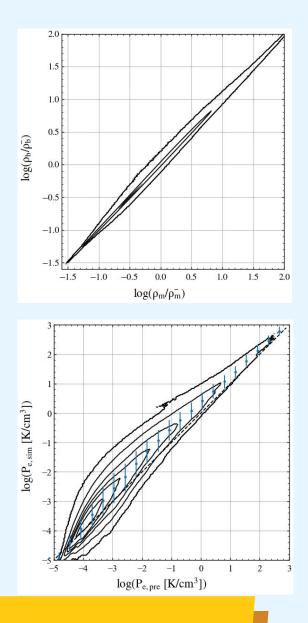




$T-\lambda$ correlation

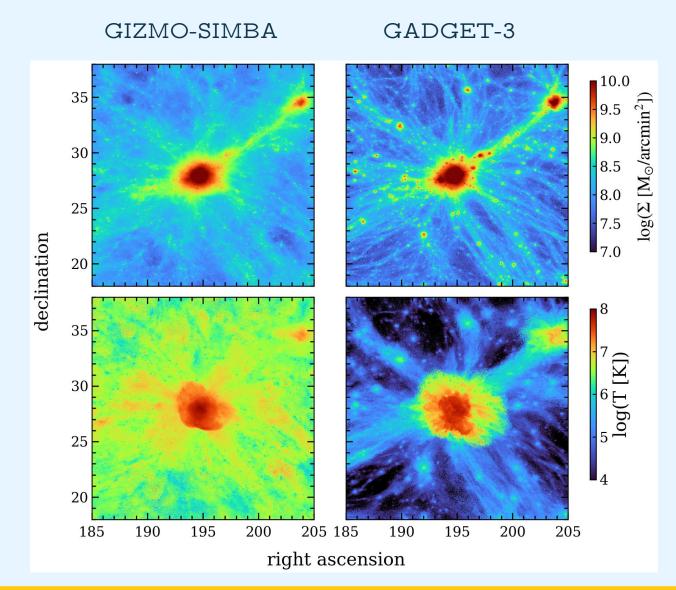
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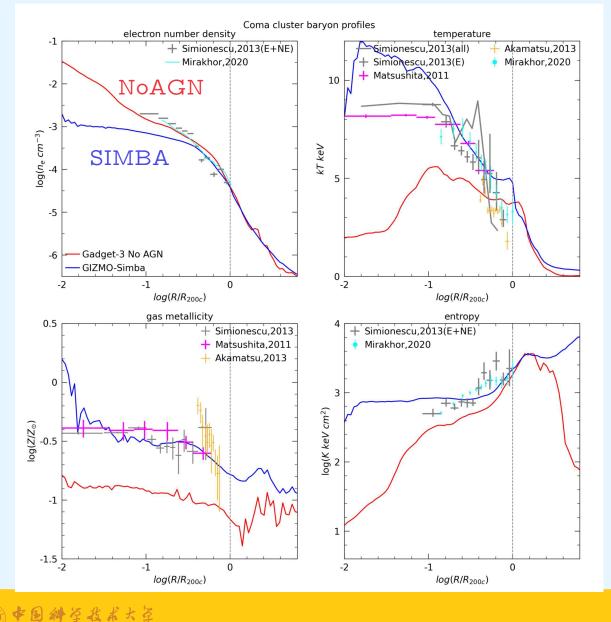
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GIZMO-SIMBA



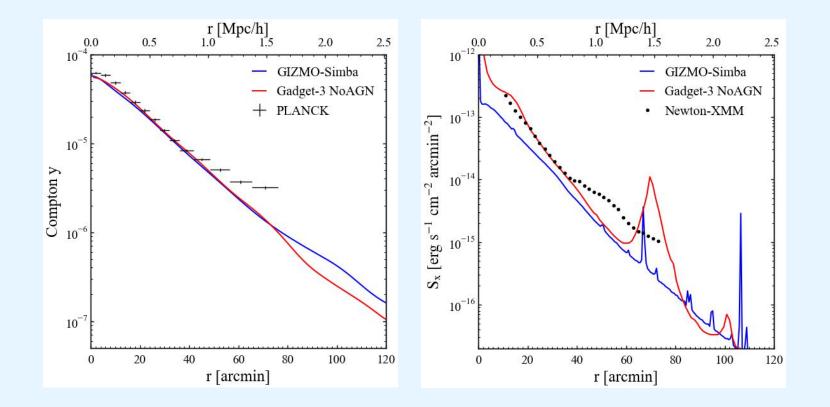
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GIZMO-SIMBA



Luo et al. in prep.

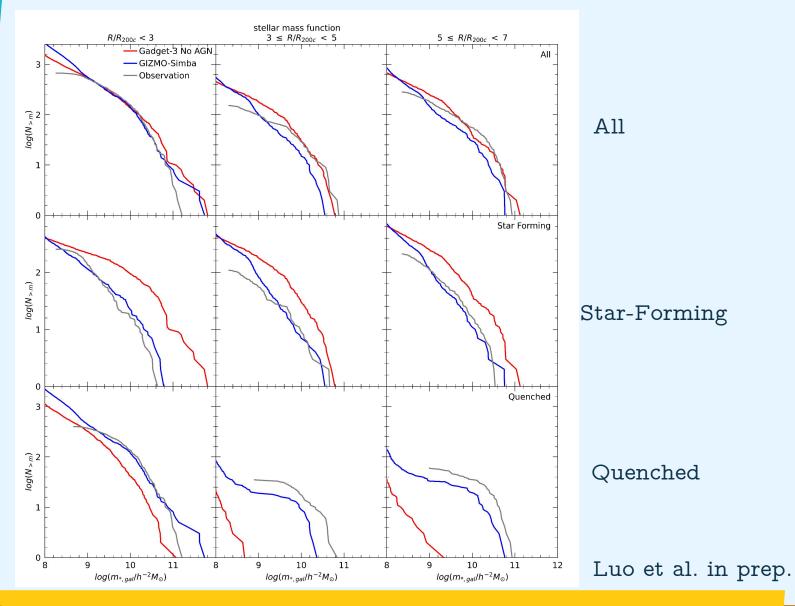
GIZMO-SIMBA



Luo et al. in prep.

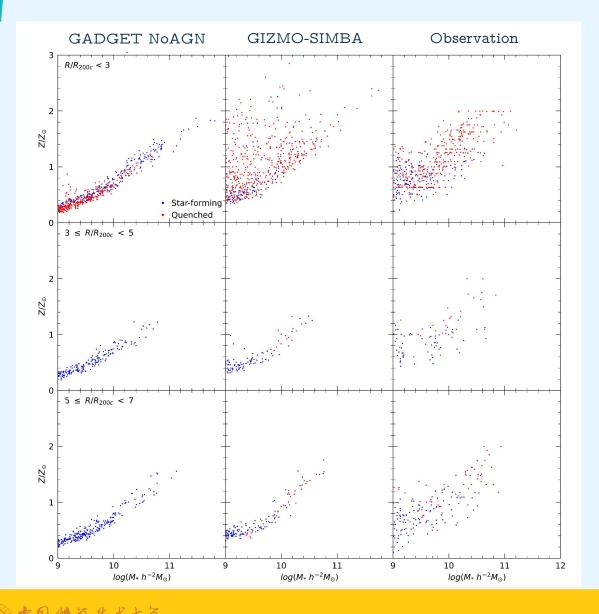
How AGN affects Coma cluster

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How AGN affects Coma cluster



Luo et al. in prep.

Thanks

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