

面向下一代大尺度巡天的超级数值模拟

- Hyper Millennium project
- PhotoNs code

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中国科学院国家天文台
NATIONAL ASTRONOMICAL OBSERVATORIES, CAS

Large simulations are still demanding for modern cosmology

- To understand complex physical processes related to various cosmic probes (BAO, RSD, Weak lensing and etc.)
- To meet the requirement of accurate Cosmology (1%)

selection effects, systematic uncertainties, statistic uncertainties

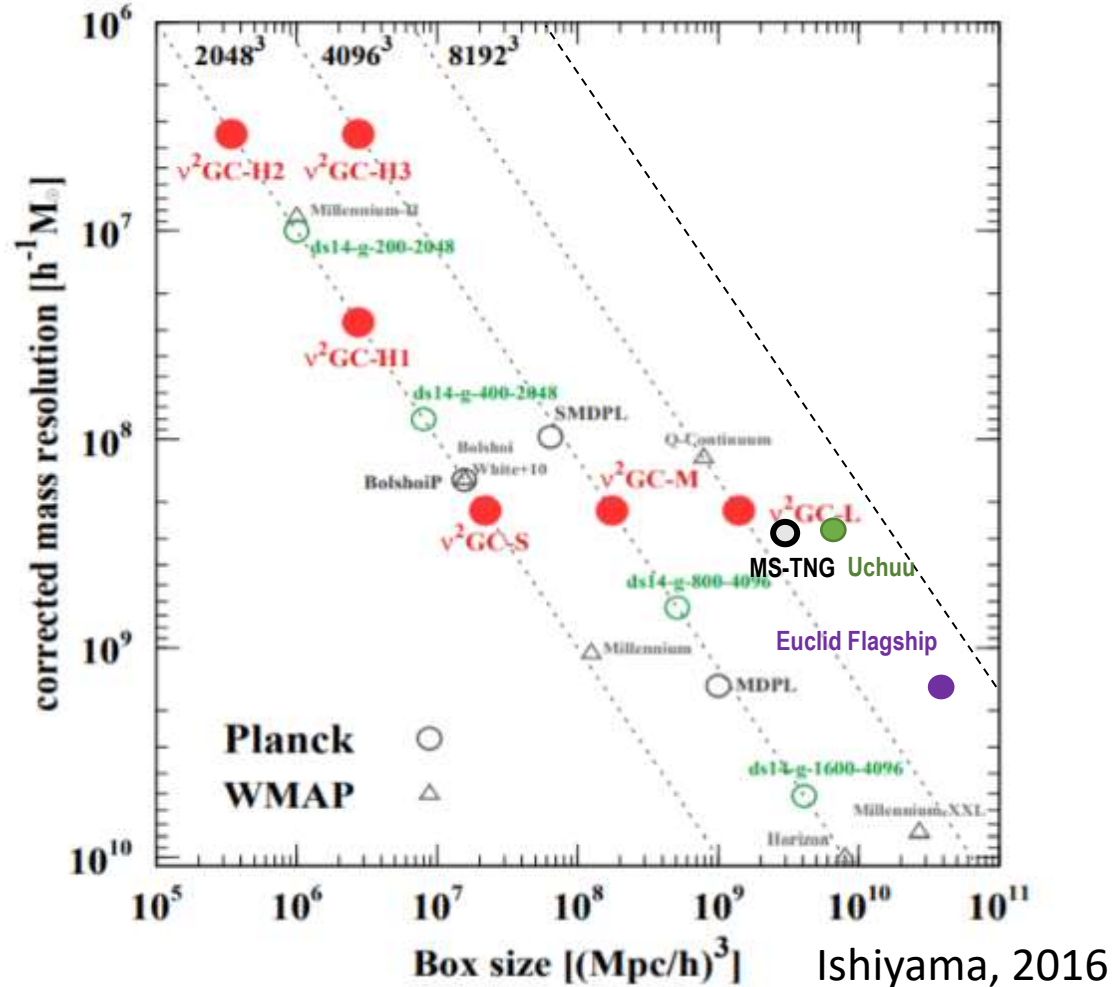
New Generation Large-scale Surveys

Name	Sky coverage (degree)	Wavelength (nm)
CSST	17,500	255-1000
LSST	18,000	320-1050
EUCLID	15,000	550-920 1000-2000
WFIRST	2,400	927-2000

Tracers:

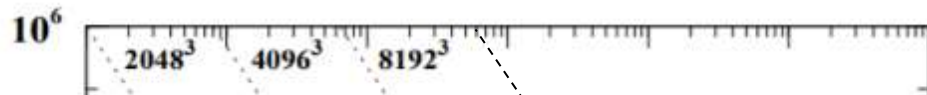
QSOs, LRGs(massive), ELGs(low masses), and etc.

art-of-the-state simulations



Name	Code	L_{box} [$h^{-1}\text{Mpc}$]	N_p [10^9]	m_p [$h^{-1}M_{\odot}$]
Horizon Run 3	GOTPM	10815	370	2.5×10^{11}
Bolshoi PL	ART	250	8.6	1.6×10^8
Small MultiDark PL	GADGET-2	400	57	1.5×10^9
MultiDark PL	GADGET-2	1000	57	1.5×10^9
Big MultiDark PL	GADGET-2	2500	57	2.4×10^{10}
Huge MultiDark PL	GADGET-2	4000	69	7.9×10^{10}
Horizon-4PI	RAMSES	2000	69	7.8×10^9
DEUS FUR	RAMSES-DEUS	21000	550	1.2×10^{12}
Pangu	L-GADGET-2	1000	29	2.5×10^9
Tiangong Pathfinder	L-GADGET-3	1000	232	3.72×10^8
$\nu^2\text{GC-L}$	GREEM	1120	550	2.2×10^8
Shin-Uchuu	GREEM	140	262	8.97×10^5
Uchuu	GREEM	2000	~ 2000	3.27×10^8
Euclid Flagship	PKDGRAV3	3780	~ 2000	$\sim 2.398 \times 10^9$
Euclid Flagship v2.0	PKDGRAV3	3600	~ 4000	$\sim 1 \times 10^9$
Q-Continuum	HACC	923	~ 550	1.05×10^8
Outer Rim	HACC	3000	~ 1070	1.85×10^9
Millennium	L-GADGET-2	500	10	8.6×10^8
Millennium-II	P-GADGET-3	100	10	6.9×10^6
Millennium-XXL	L-GADGET-3	3000	300	6.2×10^9

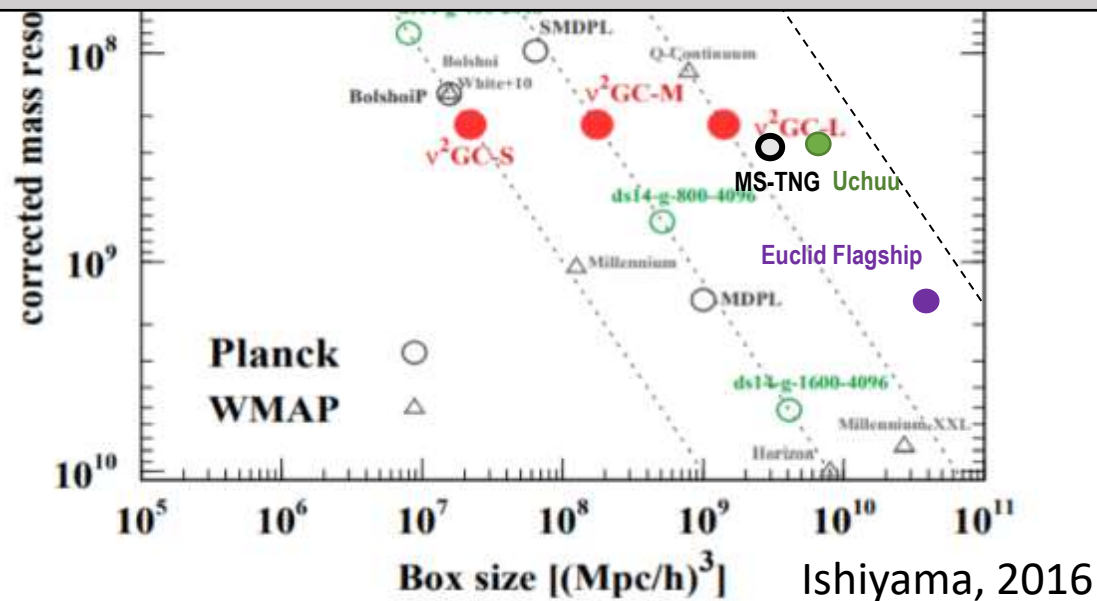
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Most current simulations can not meet the requirements

- Volume is not big enough/resolution is not high enough
- Lack of good galaxy formation models



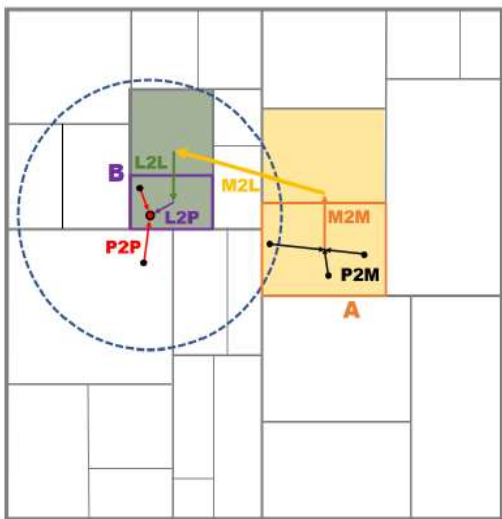
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自主研发软件平台



算法介绍

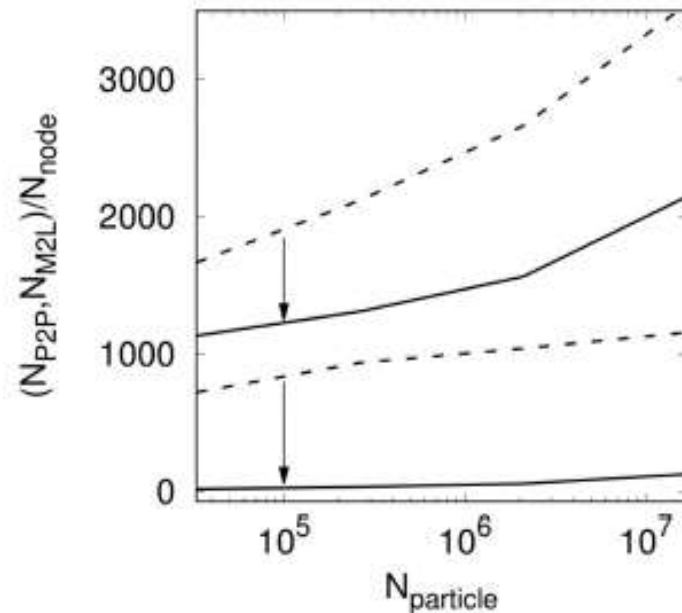
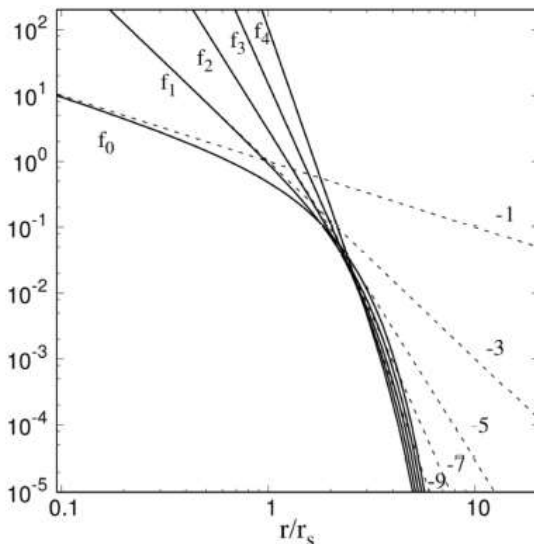
Particle-Mesh + Fast-Multipole-Method (PM-FMM)



$$f_{(2)}(r) = \frac{3}{r^5} \operatorname{erfc}\left(\frac{r}{2r_s}\right) + \frac{1}{\sqrt{\pi}} \exp\left(-\frac{r^2}{4r_s^2}\right) \times \left[\frac{3}{r_s r^4} + \frac{1}{2r_s^3 r^2} \right],$$

$$f_{(3)}(r) = -\frac{15}{r^7} \operatorname{erfc}\left(\frac{r}{2r_s}\right) - \frac{1}{\sqrt{\pi}} \exp\left(-\frac{r^2}{4r_s^2}\right) \times \left[\frac{15}{r_s r^6} + \frac{5}{2r_s^3 r^4} + \frac{1}{4r_s^5 r^2} \right],$$

$$f_{(4)}(r) = \frac{105}{r^9} \operatorname{erfc}\left(\frac{r}{2r_s}\right) + \frac{1}{\sqrt{\pi}} \exp\left(-\frac{r^2}{4r_s^2}\right) \times \left[\frac{105}{r_s r^8} + \frac{35}{2r_s^3 r^6} + \frac{7}{4r_s^5 r^4} + \frac{1}{8r_s^7 r^2} \right].$$

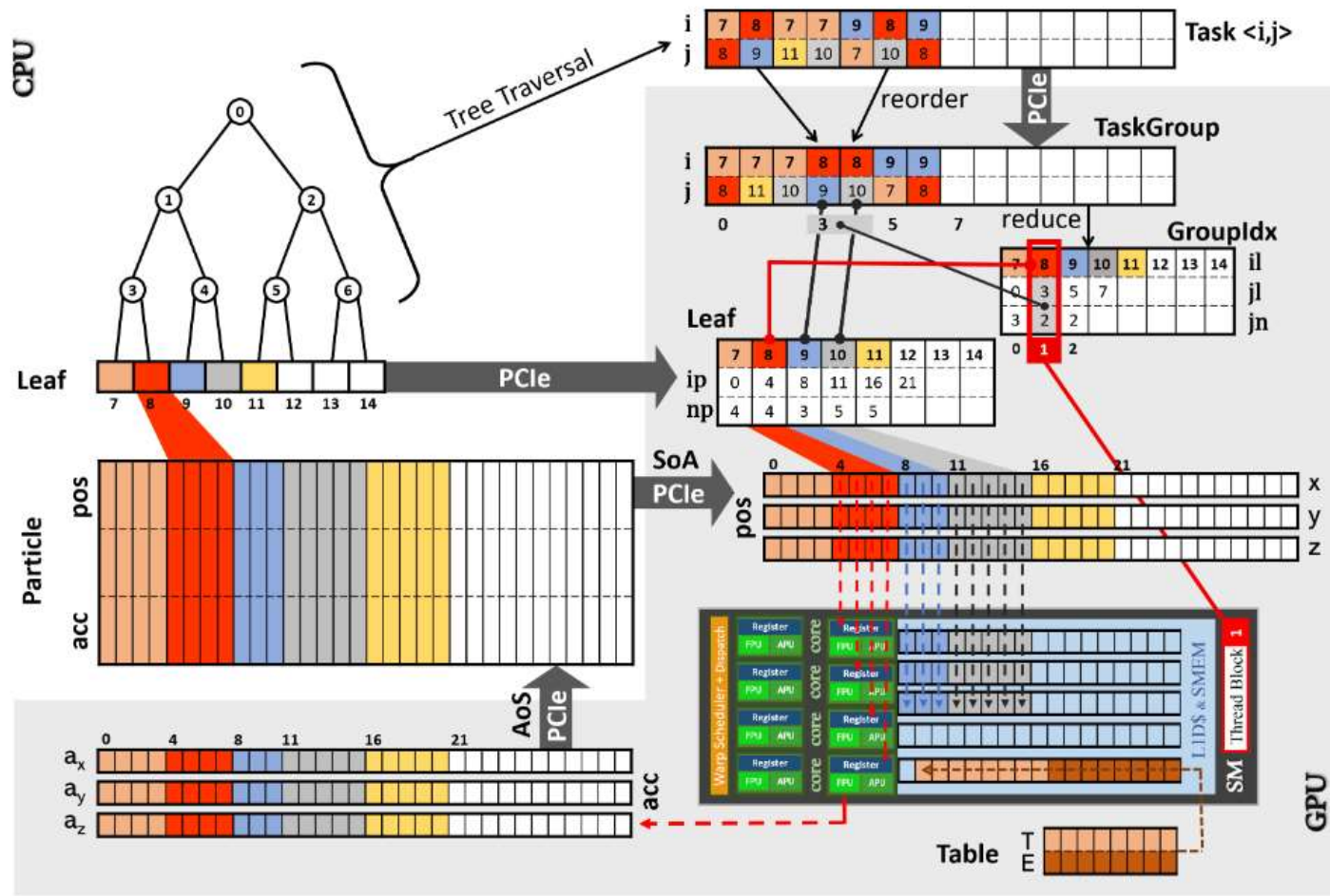


$O(N^* p + n \log_2 n + m \log_2 m)$

- PM-FMM 接近线性的时间复杂度, 同时解耦了长短程的相互作用, 极大地减少了通信和计算的总量

1. 热点计算性能和优化手段

- 树遍历的过程中通过判定条件判断建立任务列表
- 根据公式推算以及NVProf等性能分析工具得出，一次P2P核心计算，访问6个float，至少86次浮点计算
- 实现数据传输与计算重叠
- 所有源叶节点中的粒子则依次被排序处理。这种任务分配方法使得所有线程的任务完全独立，最大化访存的连续性和局部性，可以利用Global Memory的访存合并、Shared Memory共享缓存以及Register

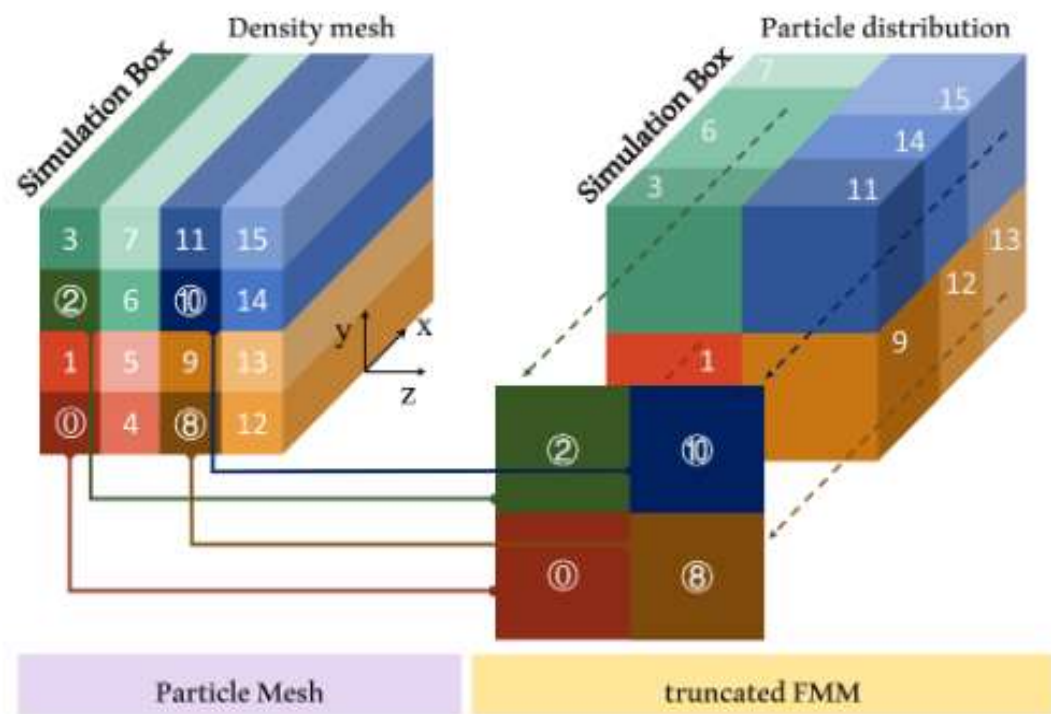


实测峰值性能 ~ 49%，平均性能 ~ 36.5% [Q.Wang & C. Meng, 2021]

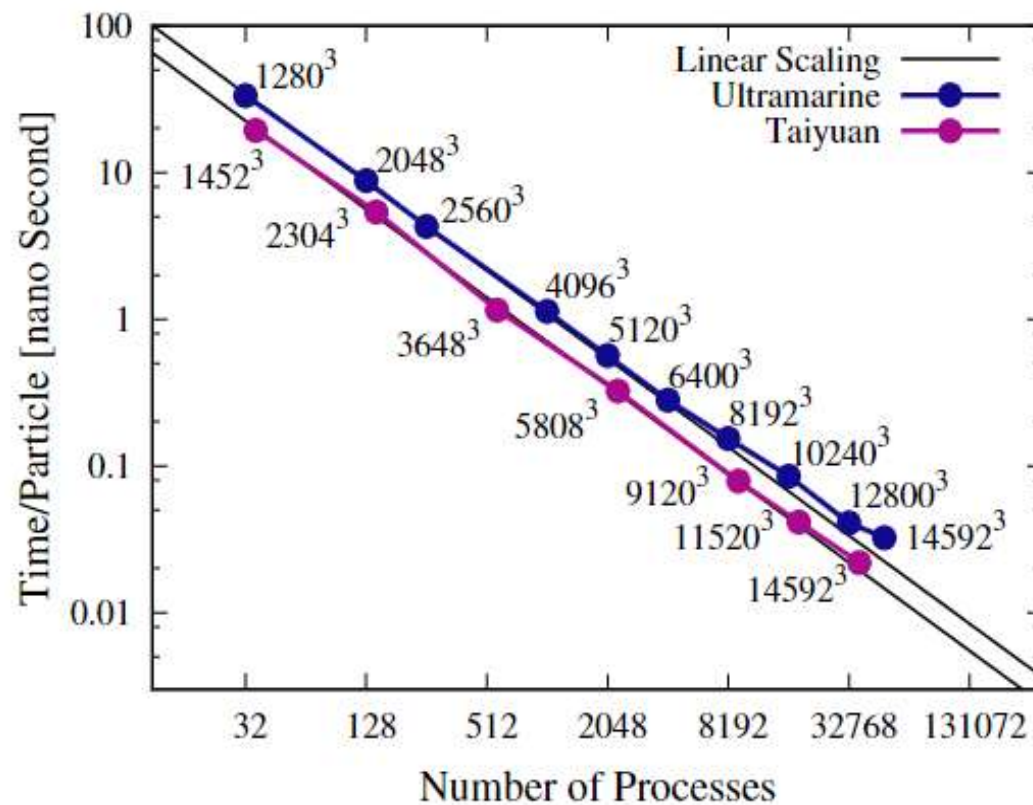
2.通信与区域分解的优化

- PM 和 FMM 的分解算法中，使用了不同的区域分解来实现
- PM 求解器执行数据和计算任务的分解
 - ✓ 对应Pencil状密度网格
 - ✓ 调用2DECOMP&FFT 库计算并执行 FFT 的全局通信
- FMM采用3D分解，优先进程组内的内存占用量均衡
 - ✓ 同组内的进程与硬件资源的绑定同样保持局部性，因为在不同的网络拓扑结构下，进程和计算资源的映射对通信性能至关重要

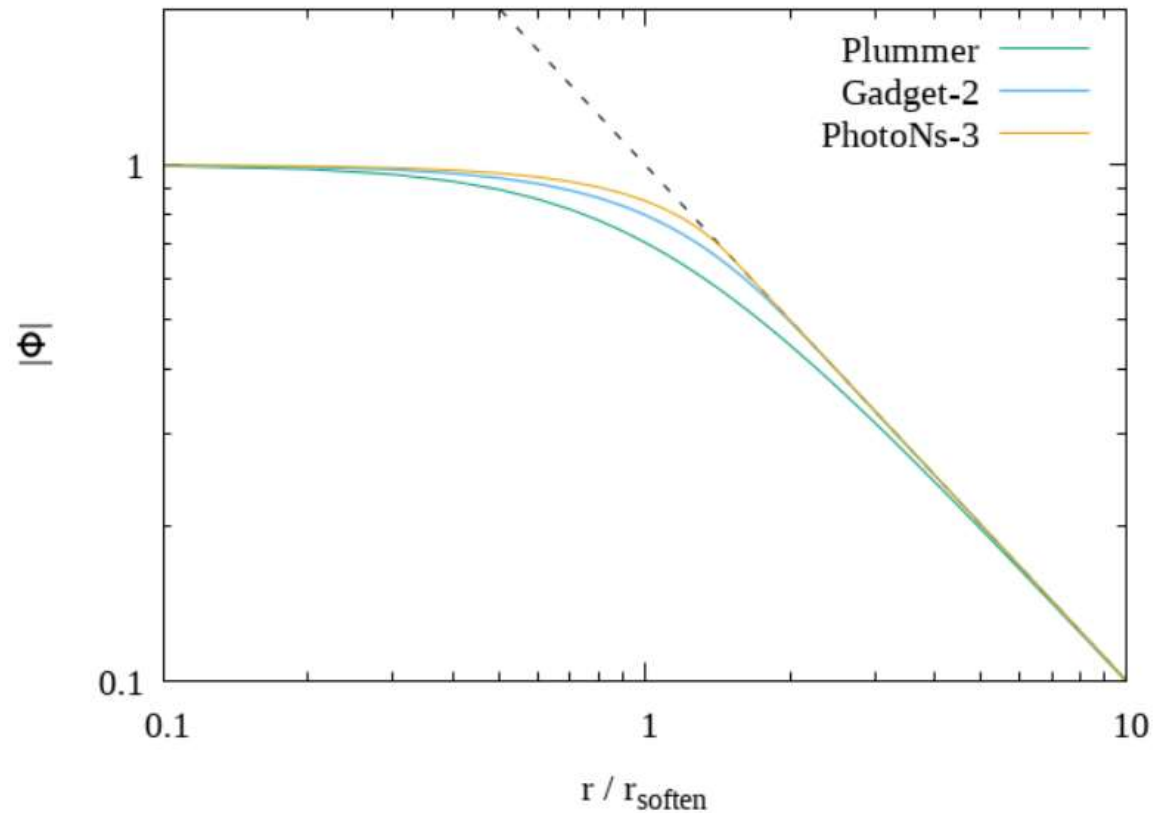
2.通信与区域分解的优化



图表 4: 并行框架与区域分解 [Wang et al. 2022 MNRAS]



Softening of Gravitational force



Gravitational potential in Plummer, Gadget-2 and PhotoNs-3, respectively.

Gravitational force goes back to inverse-square law when $r > 1.5 R_{\text{soft}}$ and it is proportional to r when $r < 1.5 R_{\text{soft}}$.

$$g(r) = \begin{cases} -G/r^2 & (r > \epsilon) \\ -Gr/\epsilon^3 & (r \leq \epsilon) \end{cases},$$

where $\epsilon = 1.5R_{\text{soft}}$.

error in PM-FMM

Multipole Acceptance Criteria

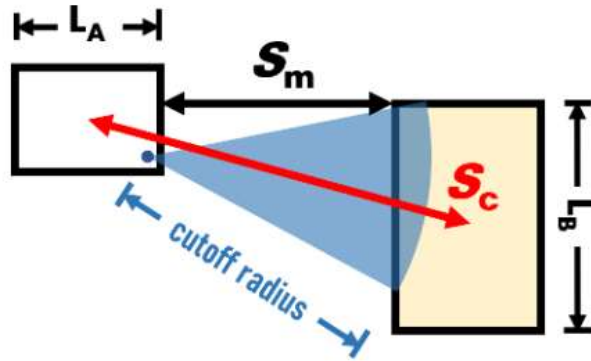


Figure 3: Multipole Acceptance Criteria. L_A is side length of sink (targeted) tree cell and L_B is the source one. The red arrow is the separation S_c between the centers of two nodes and S_m is minimum distance between two boundaries of tree cells. Two boxes are still physically relevant despite S_c beyond the cutoff radius.

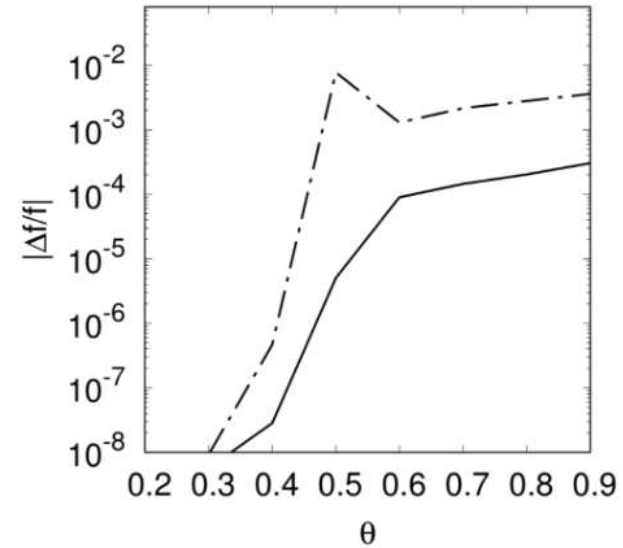
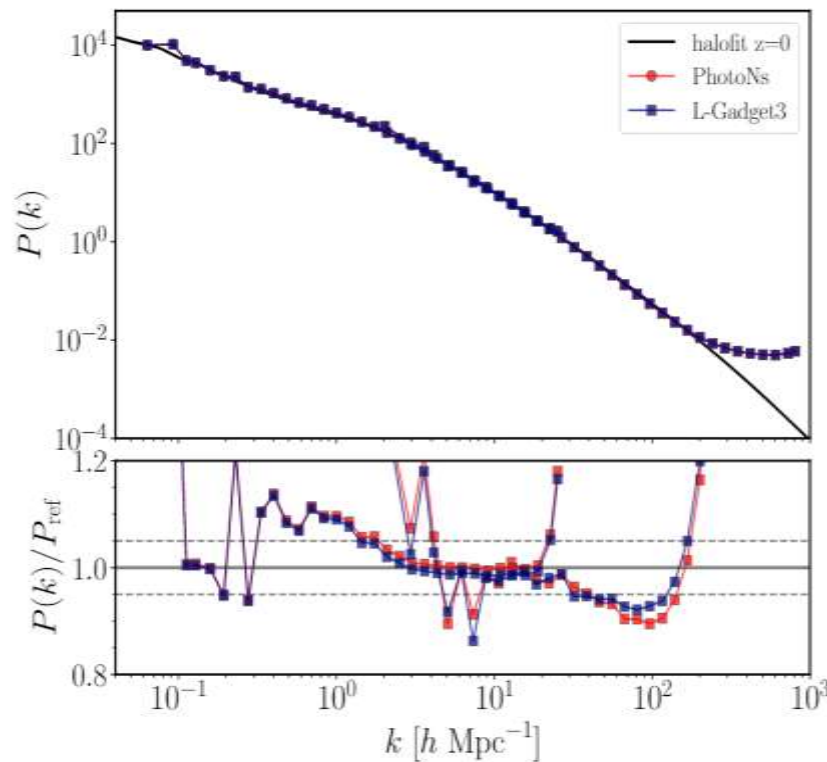


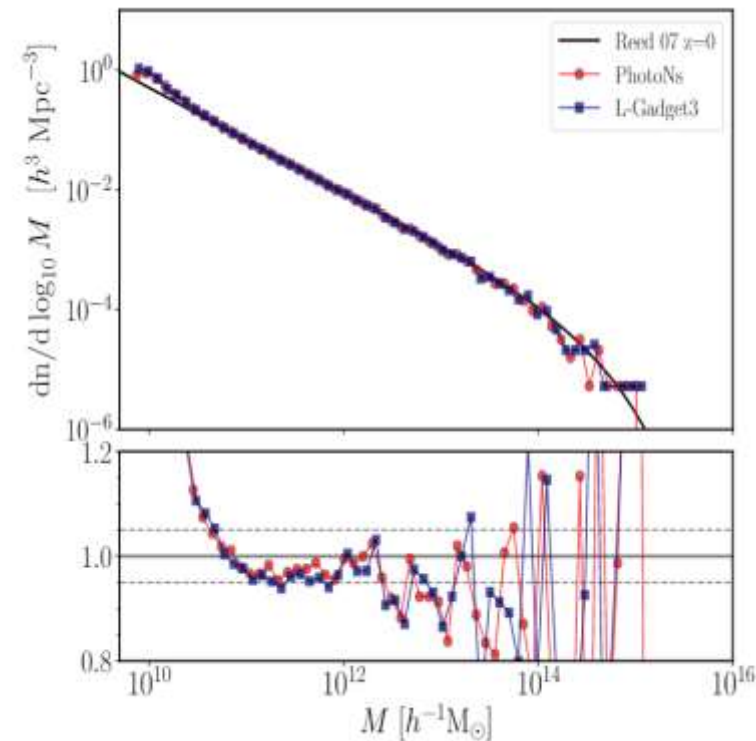
Figure 4: Relation of error to opening angle. Relative error of gravity or acceleration is calculated via truncated FMM upto hexpole. The solid curve denotes the rms error and dash-dotted curve denotes the maximum error of all particles.

与L-Gadget3比较测试PhotoNs-3的计算精度和数值收敛性

- 使用相同的IC同时运行
768³, box=120Mpc/h模拟,
比较两点相关和物质功率谱
- 不同程序之间P3/G3, 在大尺度 ($k \leq 1$) 上 $<2-3\%$ 的差别, 在 $k \leq 10$ 的尺度上 $<1\%$, 与Gadget4的比较有类似的结构。



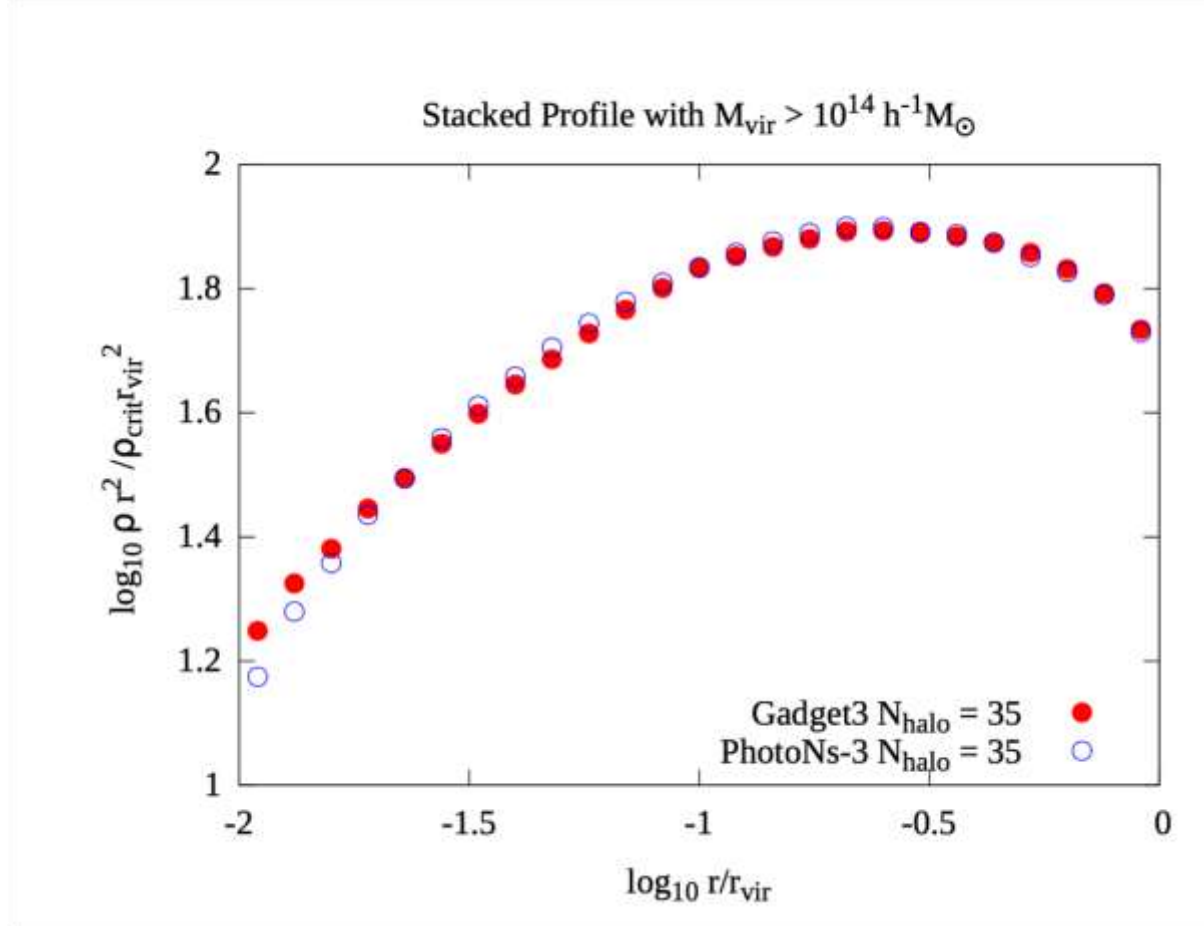
物质功率谱



暗晕质量函数

Comparison between PhotoNs and Gadget

Stacked cluster density profiles

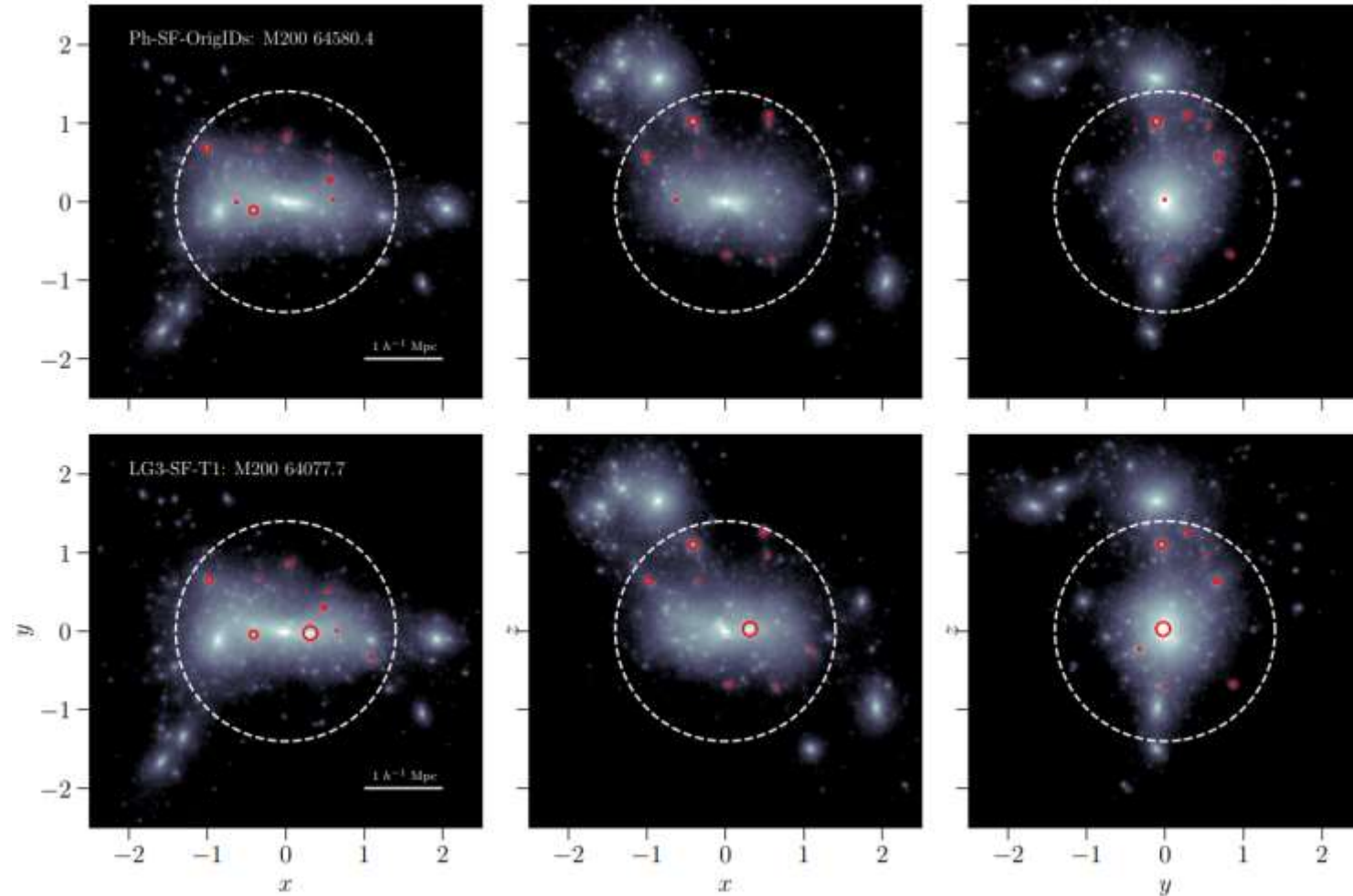


768^3 particles + 125Mpc/h

Comparison between PhotoNs and Gadget

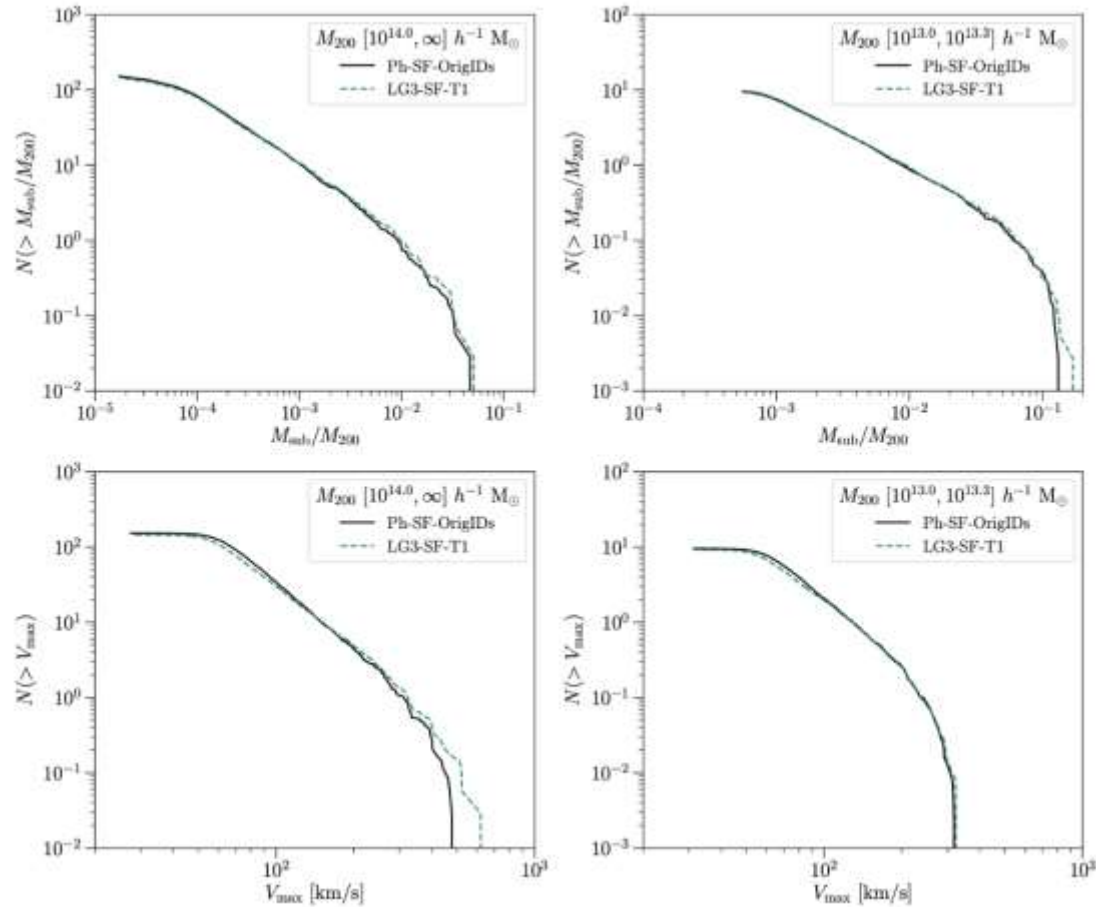
Projections of top 10 most massive subhalo in the most mass cluster

768^3 particles + 125Mpc/h



Comparison between PhotoNs and Gadget

Subhalo mass functions



768³ particles + 125Mpc/h

Hyper-Millennium run

Large volume + high resolution ($z = 20 - 0$, 100 snapshots)

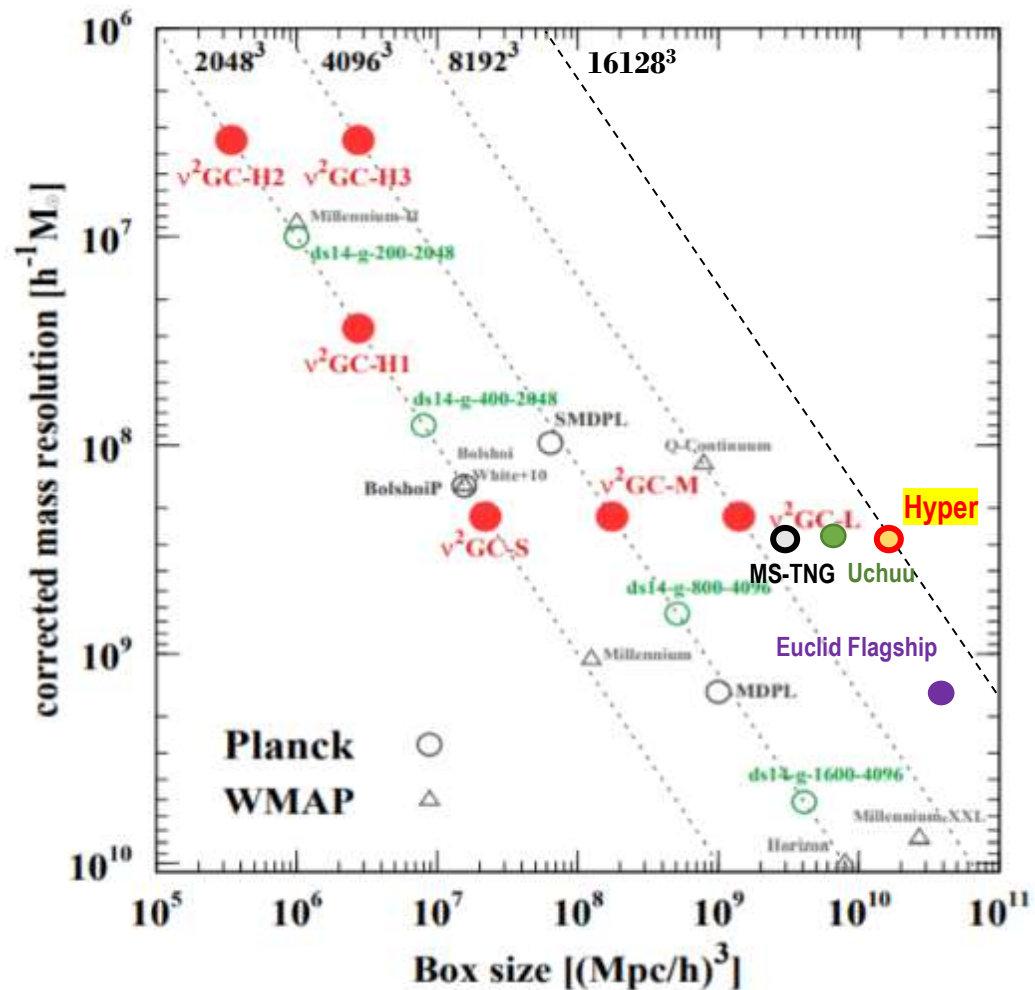
Ω_m	0.3111
Ω_Λ	0.6889
Ω_b	0.0490
h	0.6766
σ_8	0.8102
n_s	0.9665

	Hyper-Millennium
Volume	$2.5^3 [h^{-3} \text{Gpc}^3]$
Particles number	$16128^3 \sim 4.195 \times 10^{12}$
Particle mass	$3.21 \times 10^8 h^{-1} M_\odot$

Largest ever N-body cosmological simulation

Volume is 125 times the Millennium simulation and the resolution is higher by a factor of 3

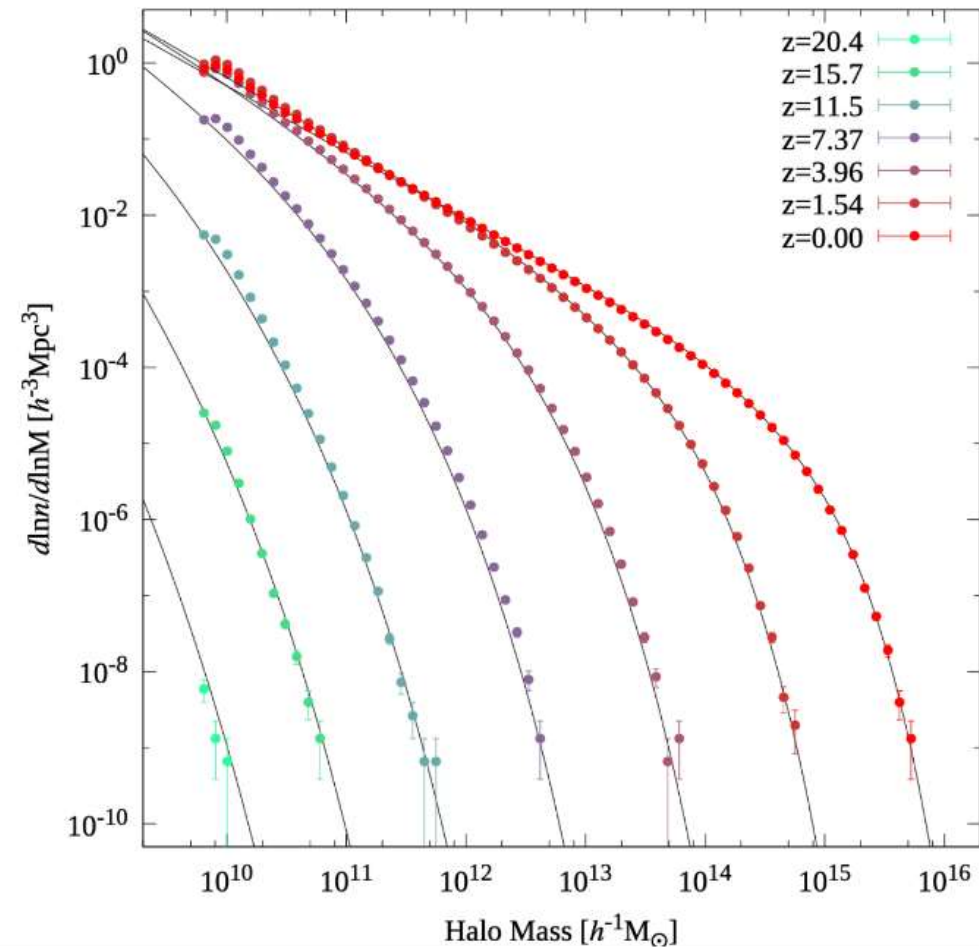
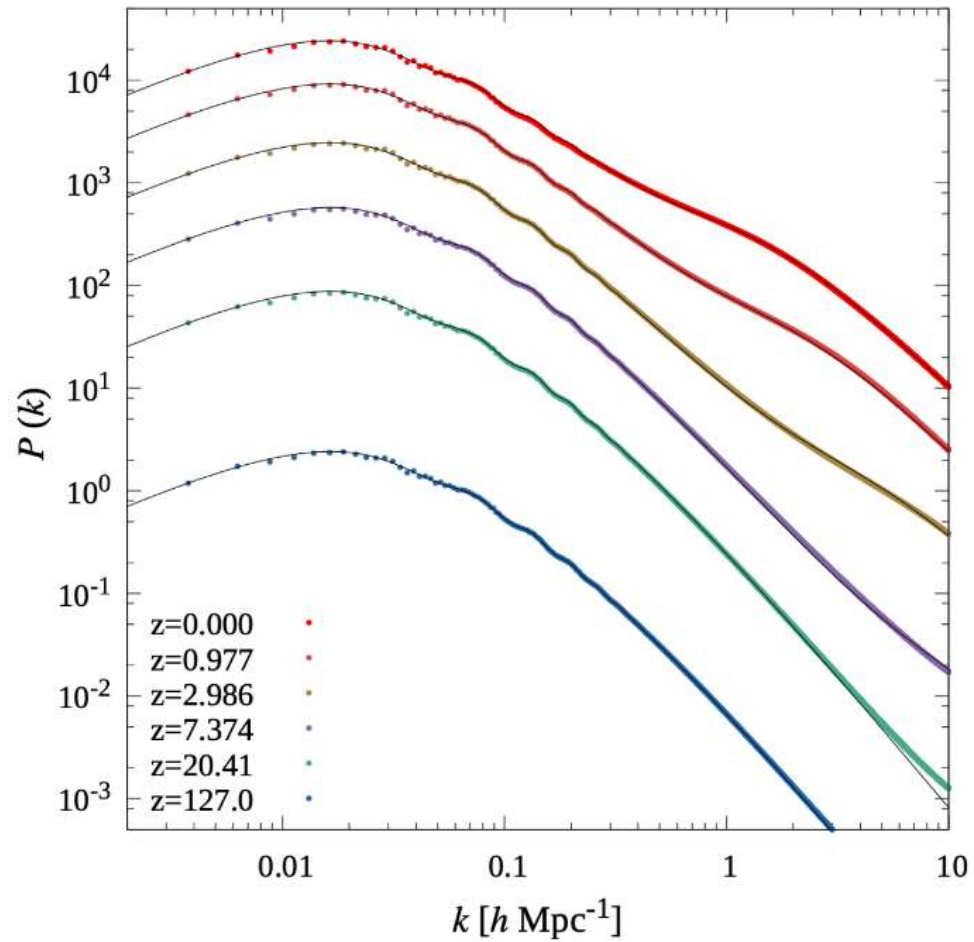
Hyper-Millennium run



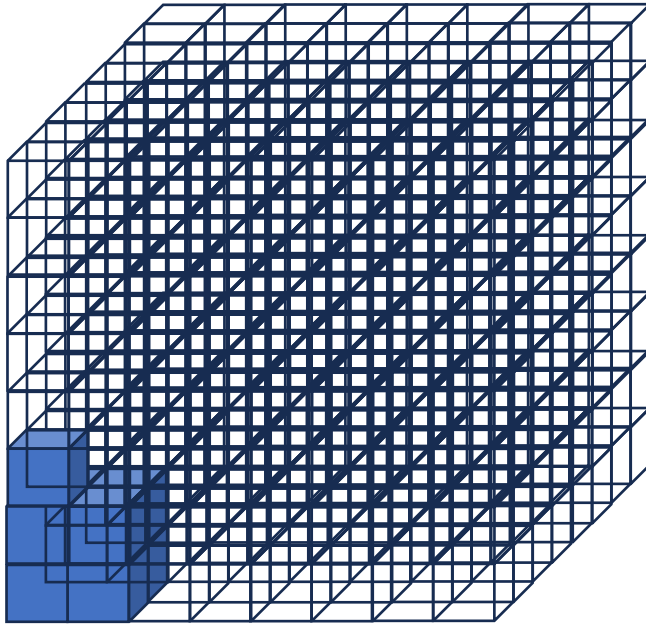
- 东方一号运行墙钟时间17.5天，13 PB的输出



Hyper-Millennium: Power spectrum and the halo mass functions



Distributed data Post-processing

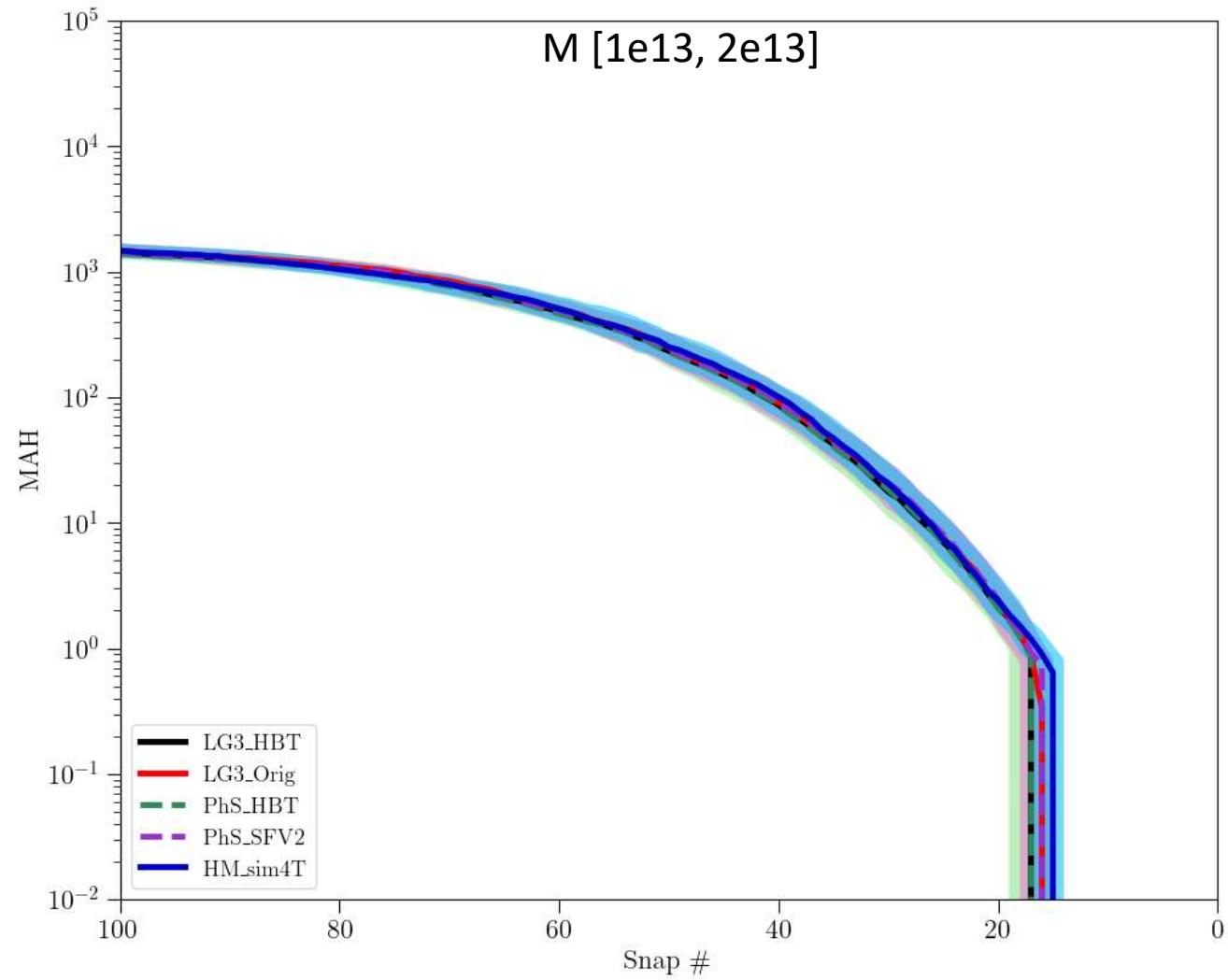


Codes validations are nearly finished

Massive postprocessing will start very soon and should be completed in this August.

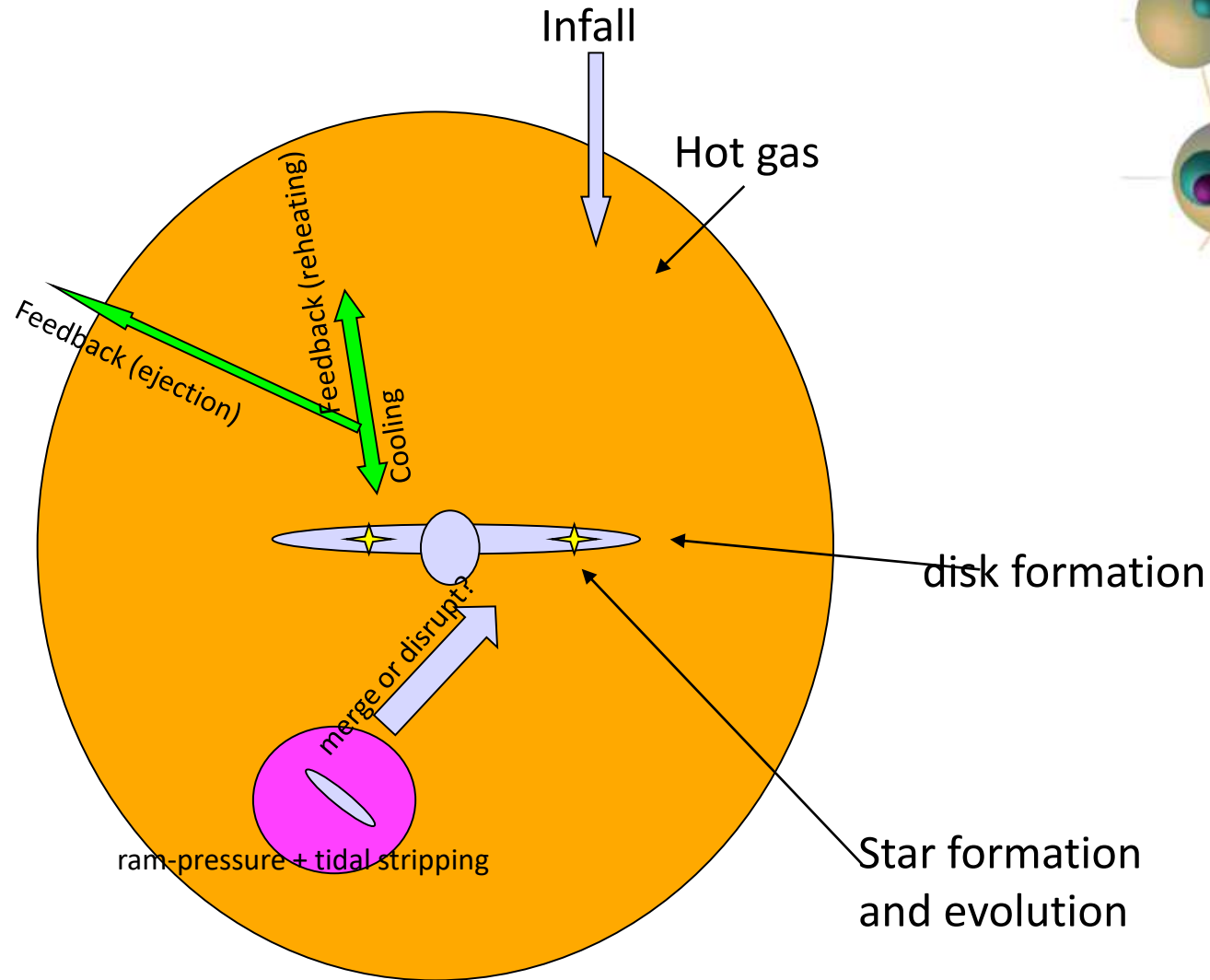
- Halo/subhalo are identified in $8 \times 8 \times 8 = 512$ cubes (330Mpc/h on a side)
- Only inner 300Mpc/h are used. The cubes are combined later to obtain the complete catalogues/merging trees in the whole volume
- In each cube, **~1TB** for halo+subhalo catalogues in format of hdf5. **~500TB** in **total**.
- MPA-tree ~50 TB
Aux file ~ 350 TB

Halo Mass Accretion History

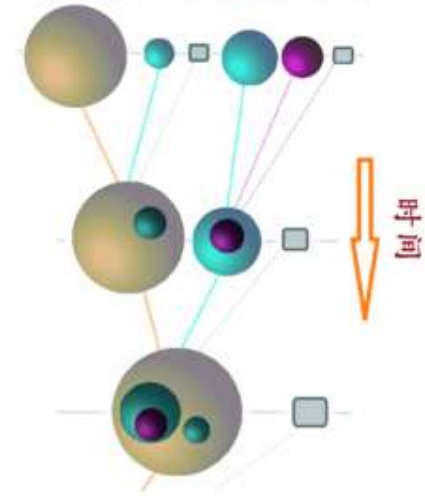


Galaxy formation models

<http://galformod.mpa-garching.mpg.de/mrobs/private/china.jsp>



暗物质晕的逐级形成过程

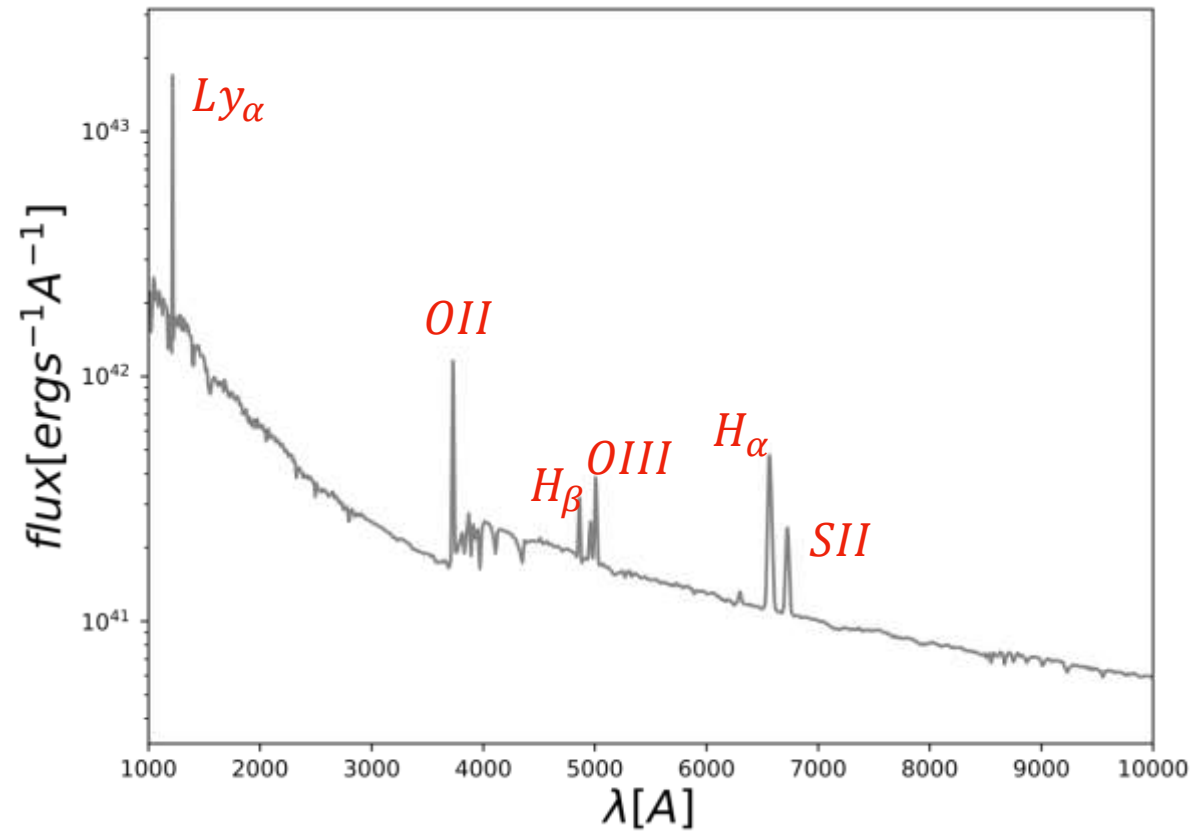


- Emission lines
 - Line ratios - CLOUDY13.03
 - geometry
 - chemical content
 - ionizing spectrum
 - metallicity (Z)
 - ionizing parameter (U)
 - Hydrogen density (n_H)

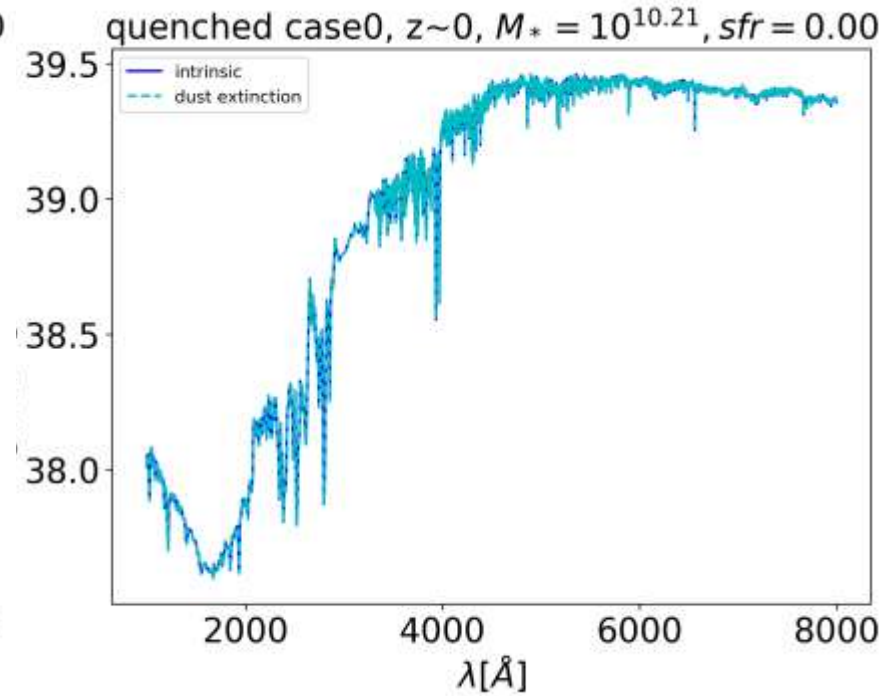
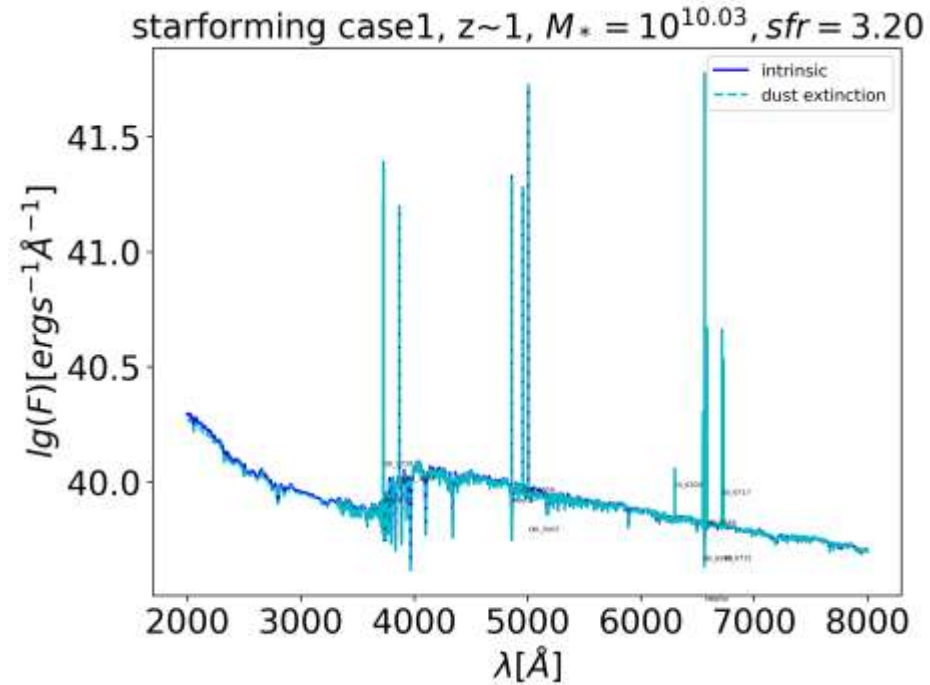


Wenxian Pei

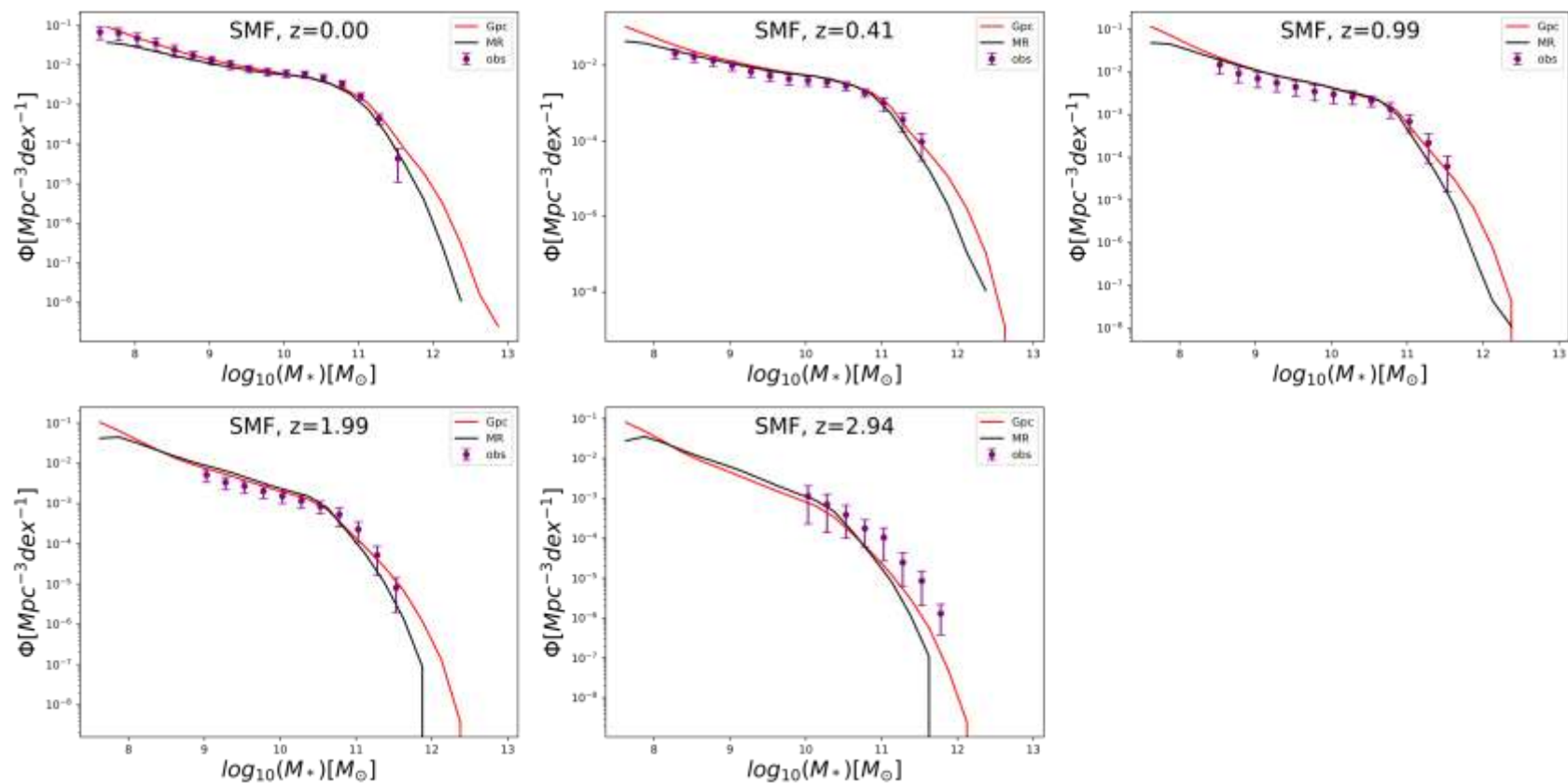
star forming galaxies at $z \sim 0$



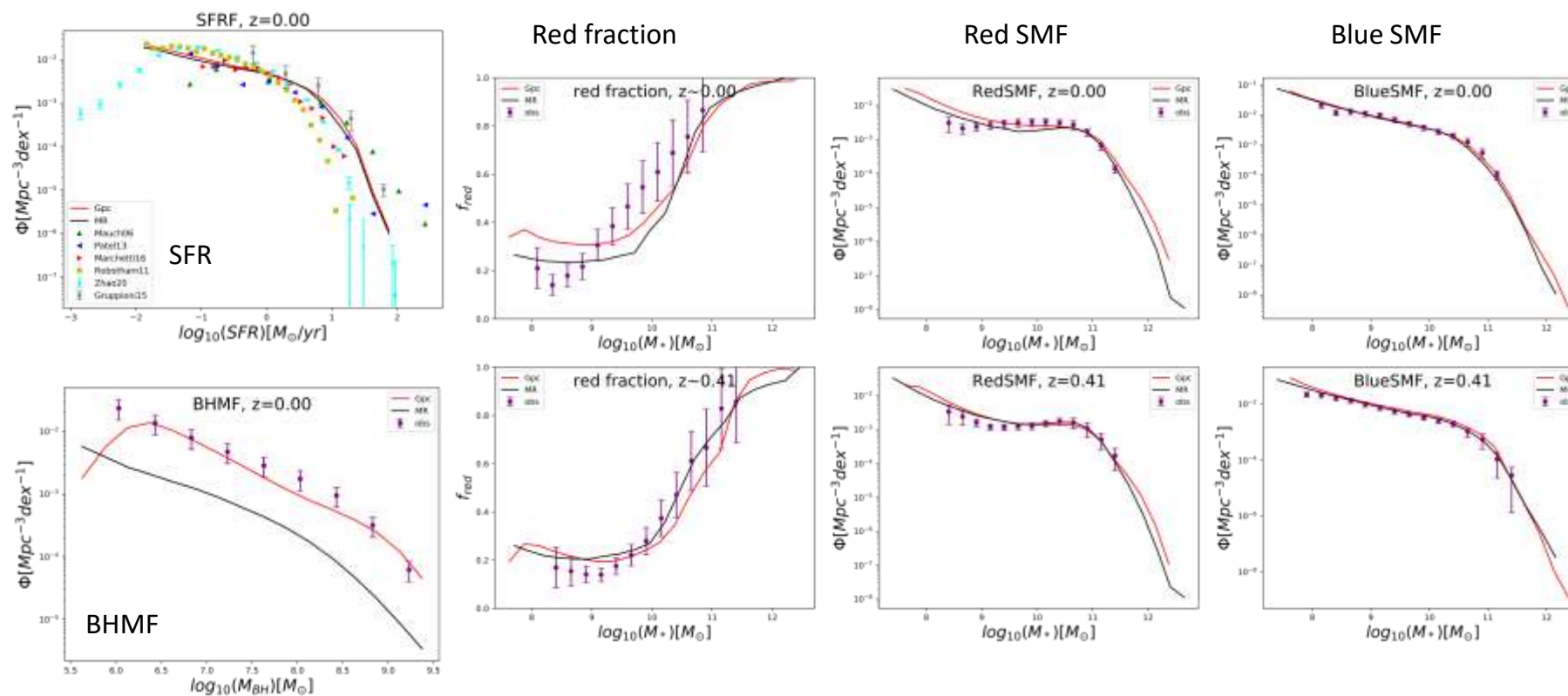
Full Galaxy Spectrum



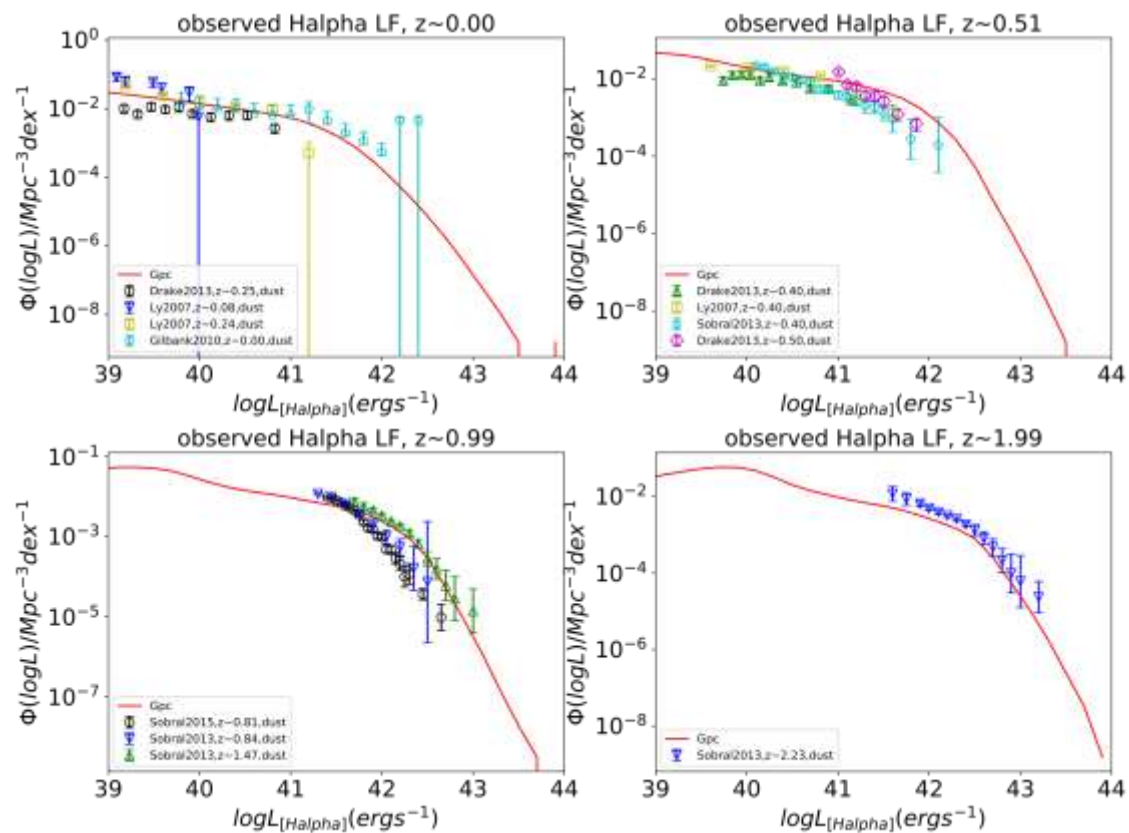
Reproduce stellar mass functions $z=0-3$



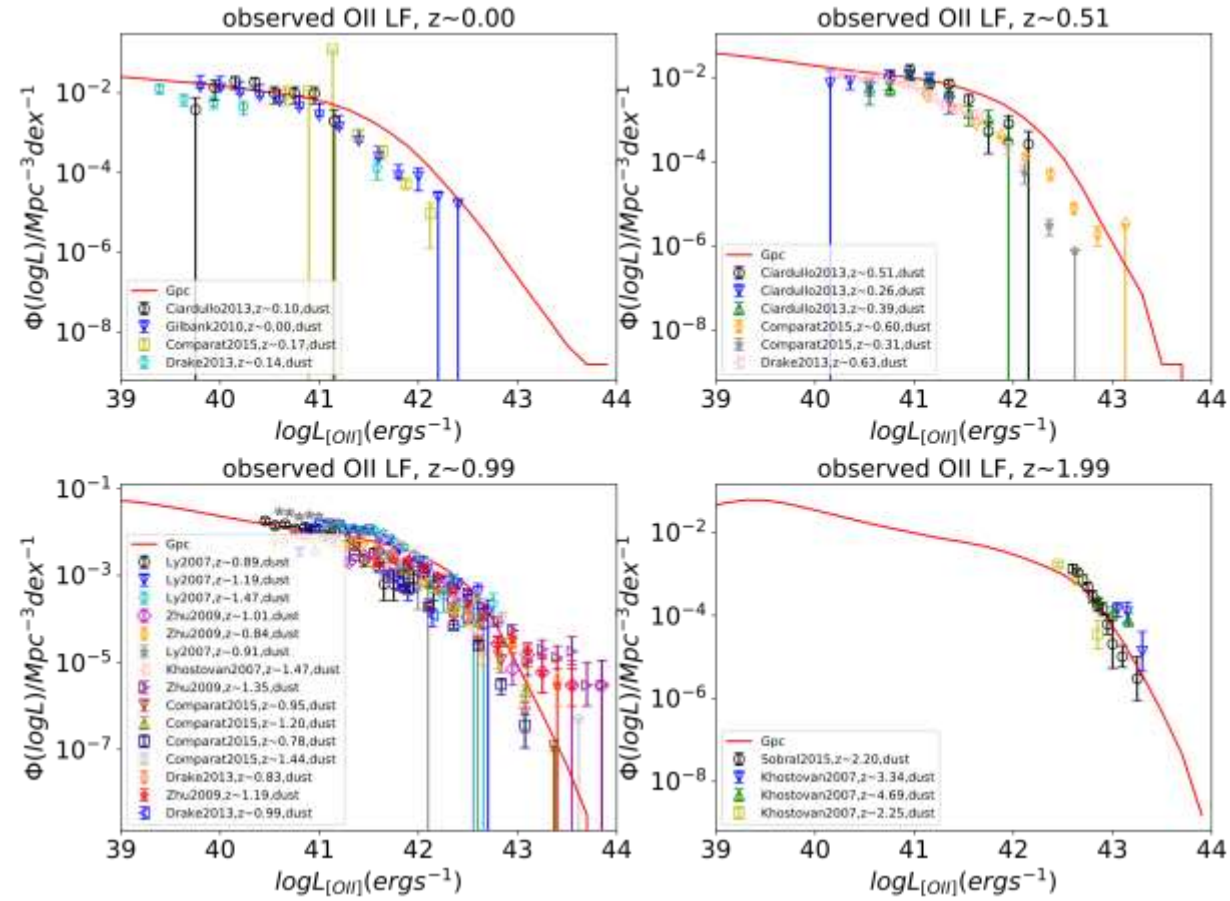
Reproduce various galaxy properties



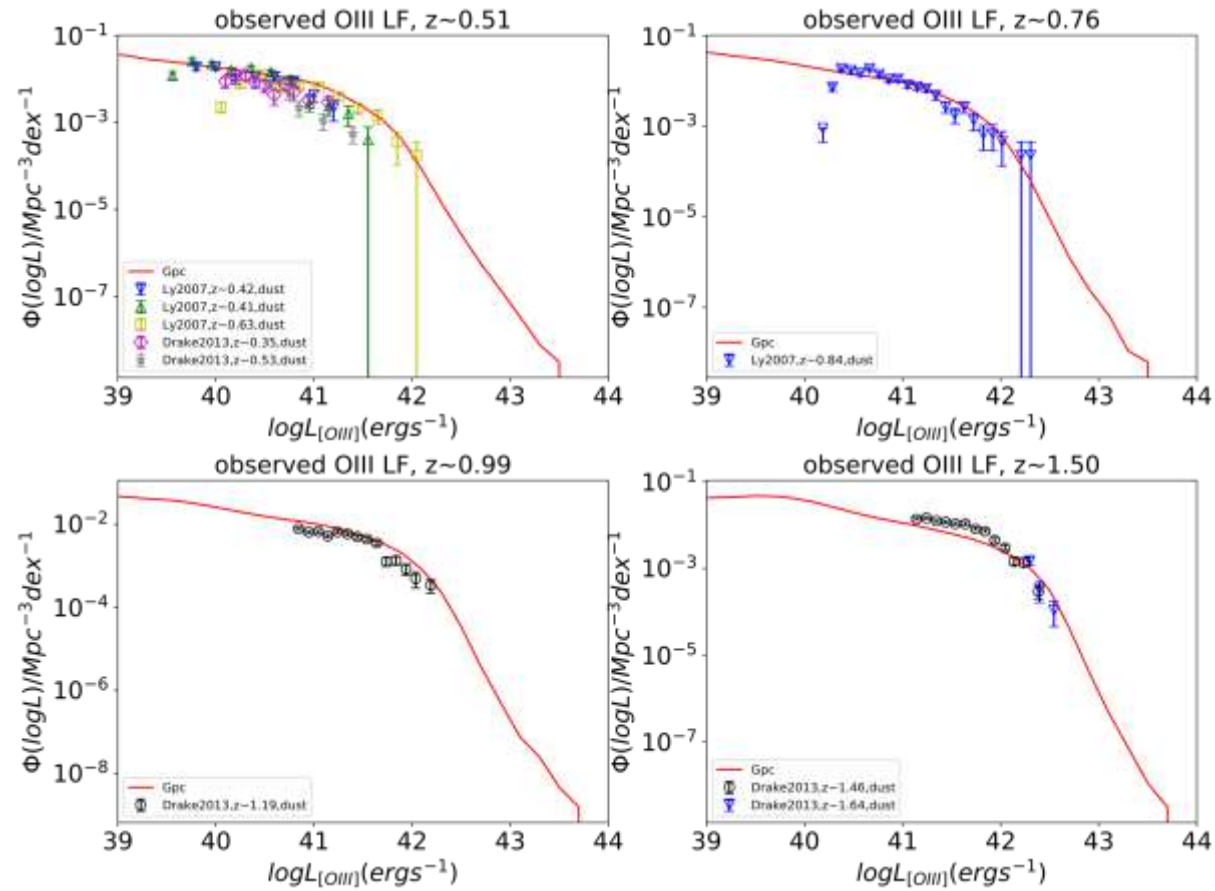
H_{α} Luminosity Function $z=0-3$



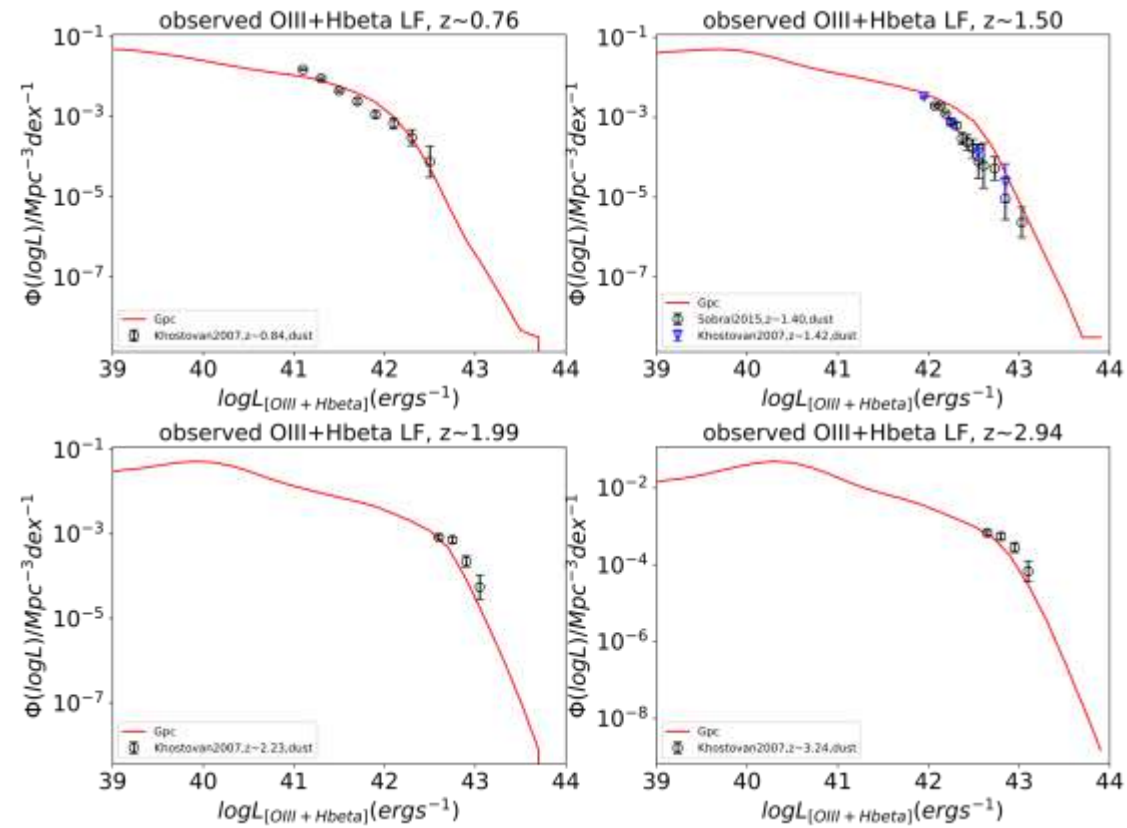
OII Luminosity Function $z=0-3$



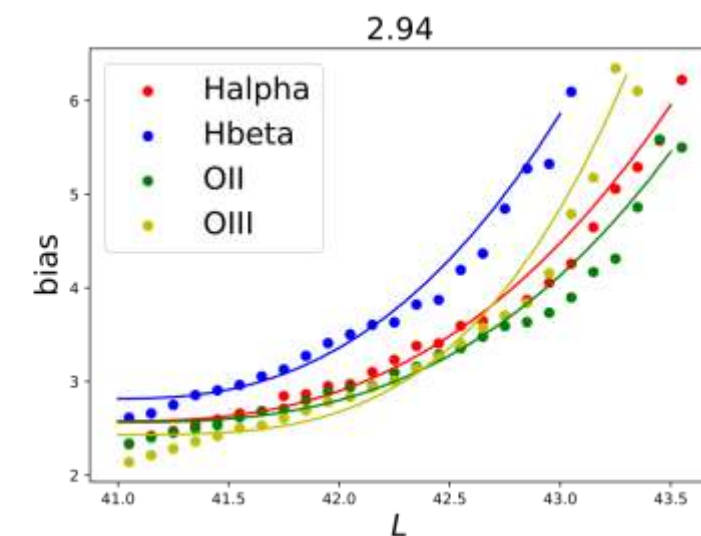
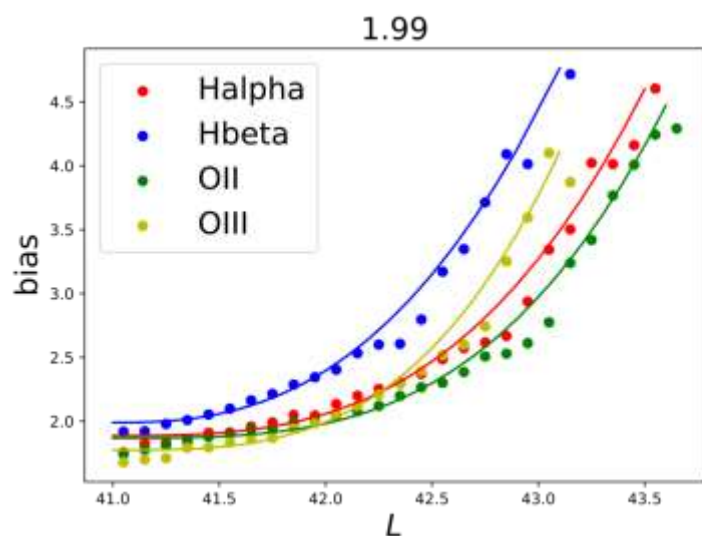
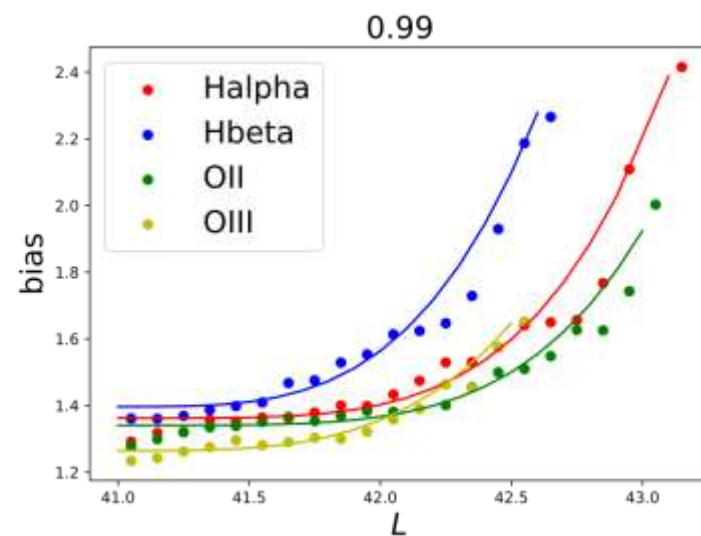
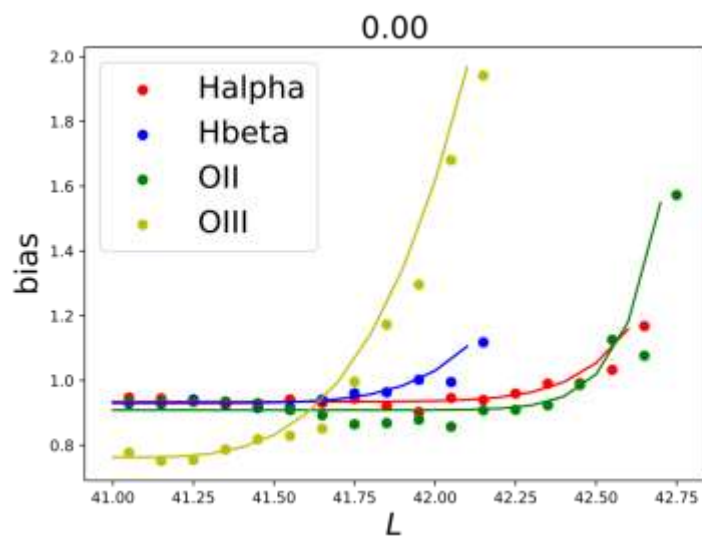
OIII Luminosity Function $z=0-3$



$OIII + H\beta$ Luminosity Function $z=0-3$



Bias as a function of luminosity and z



Summary

- Hyper—Millennium run--the largest ever N-body simulation
4 trillion particles, $2.5^3[h^{-3}\text{Gpc}^3]$, particle mass $3.2d8\text{Msun}/h$
- Improved Munich semi-analytical galaxy formation models
 - Convergence issue fixed
 - Ionization model + radiation transfer (galaxy emission lines)
 - Succeed in reproducing local and high redshift galaxy properties, especially including emission line luminosity functions of H α , H β , OII, OIII, and AGN luminosities
- To come...
 - Halo/subhalo catalogues, merging trees (before this September)
 - Mocks (galaxy, Lensing) ...
- Science
BAO, RSD, Weaking lensing and other large scale structure studies....



谢谢！ 请批评指正！



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