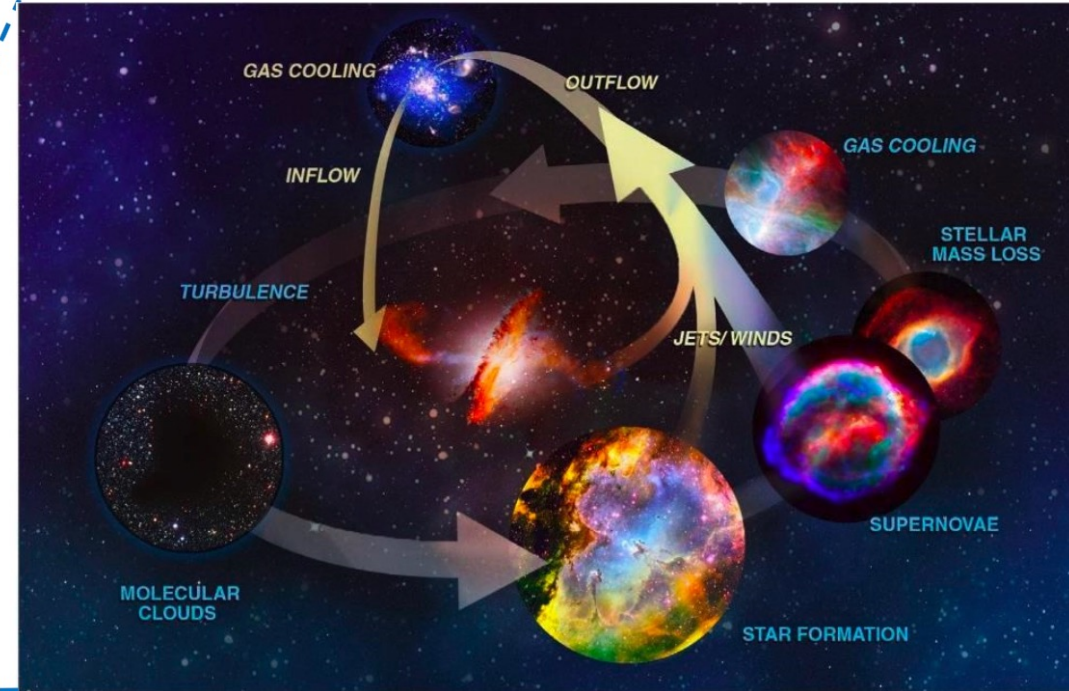
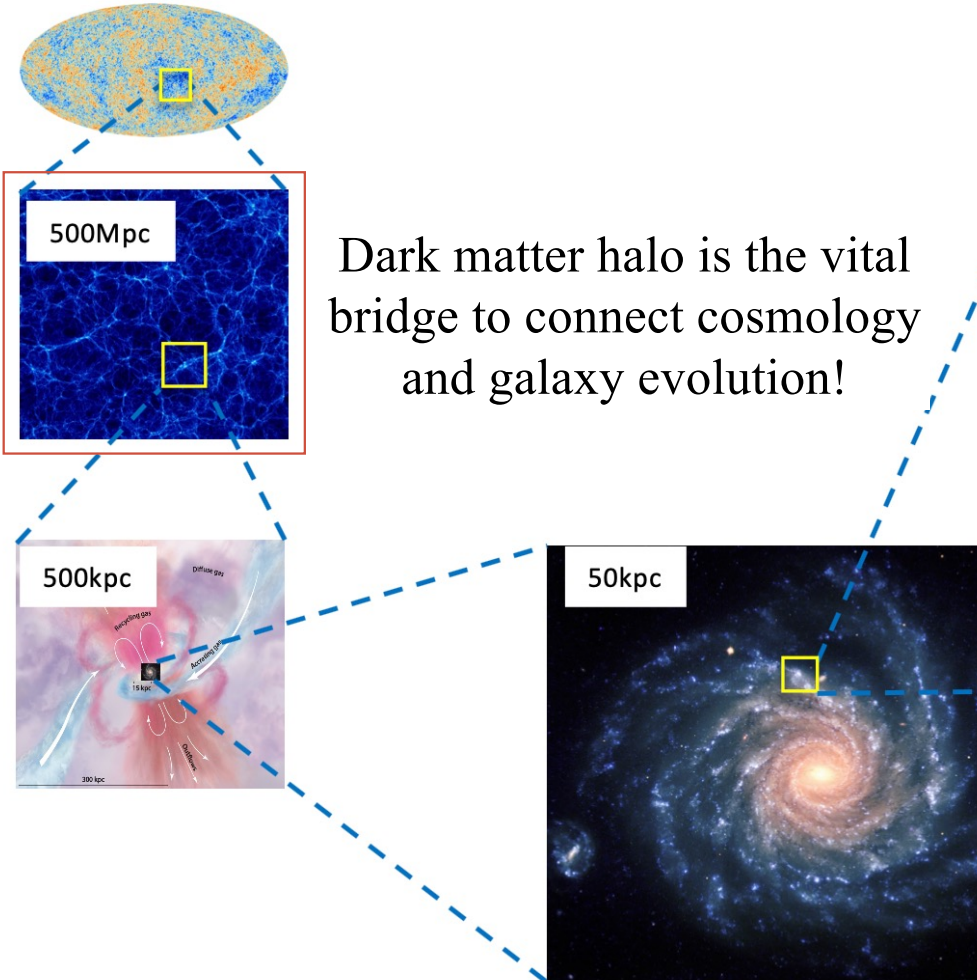


# Dependence of the dark matter halo assembly bias on halo definition and orientation

Qinglin Ma (马庆麟),  
Collaborator: Cheng Li

Oct30 - Nov 3, The 2nd Shanghai Assembly on Cosmology and Structure Formation

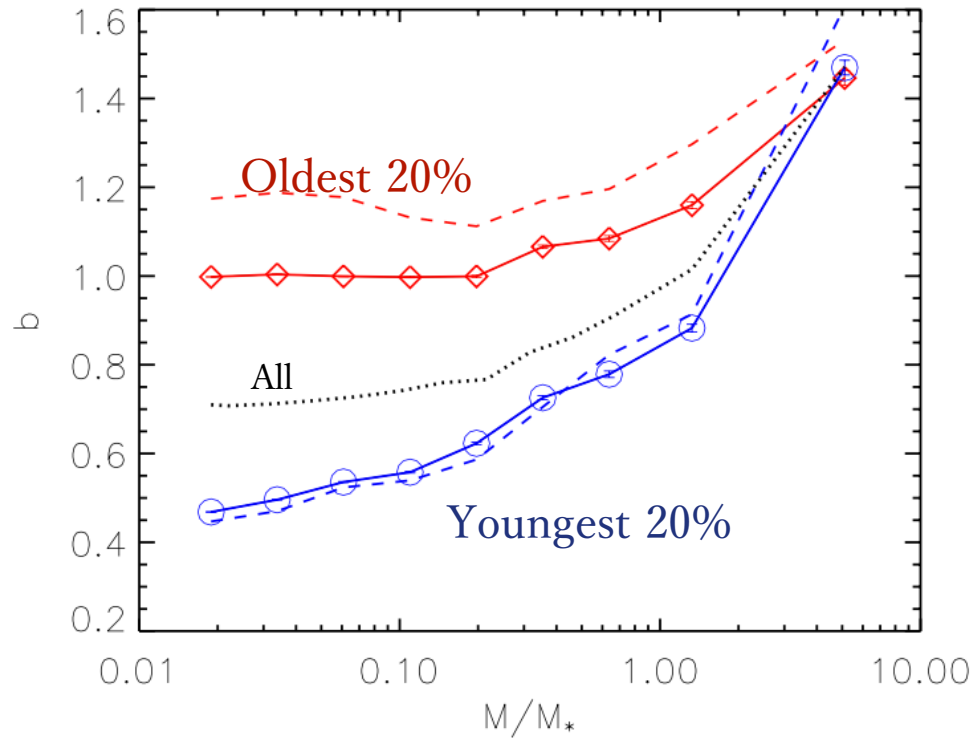
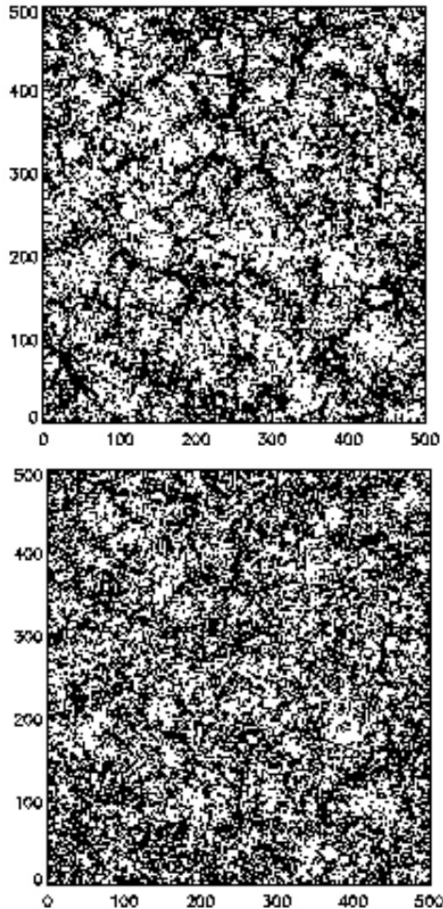
# Galaxies form and evolve in dark matter haloes



(see Astro2020)

# The halo bias is not purely determined by the halo mass.

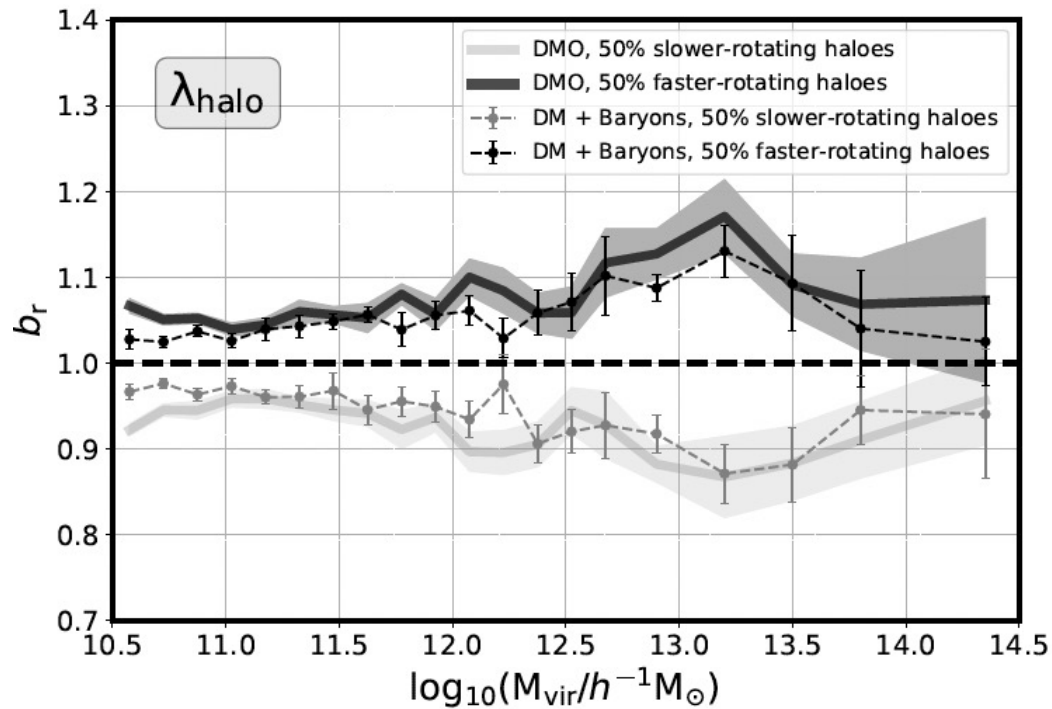
**Halo assembly bias:** Old haloes are strongly clustered than young haloes



*Gao et.al 2005*

The halo assembly bias is related to halo properties, such as formation time, concentration, shape, spin ...

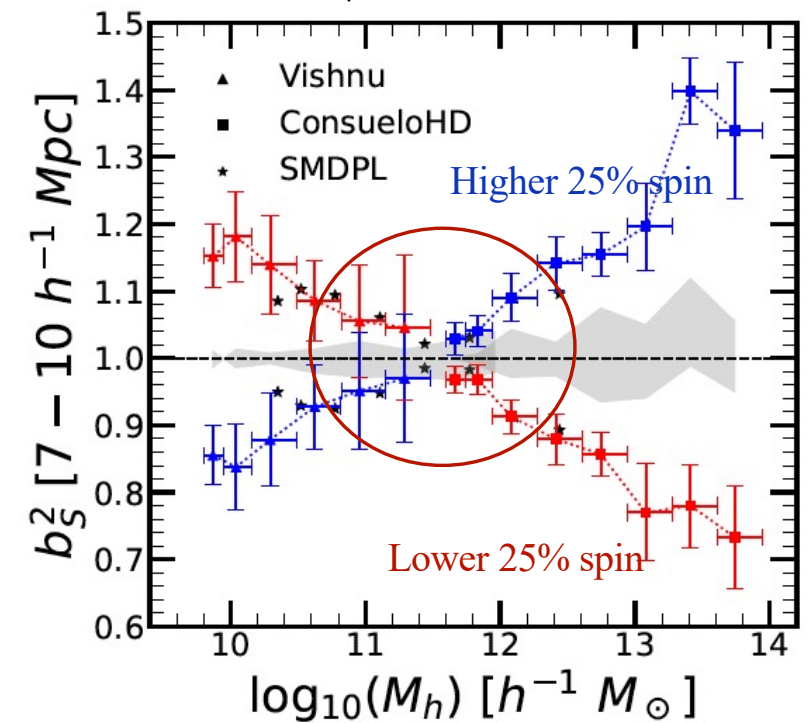
# A tension of the spin bias



Montero-Dorta et.al 2021

$$b_{\lambda}^{\text{rel}}(r) = \sqrt{\frac{\xi_{hh}(r|M_h, S)}{\xi_{hh}(r|M_h)}}$$

Tension!



Johnson et.al 2019

# The different trend of the spin bias is caused by different halo definitions

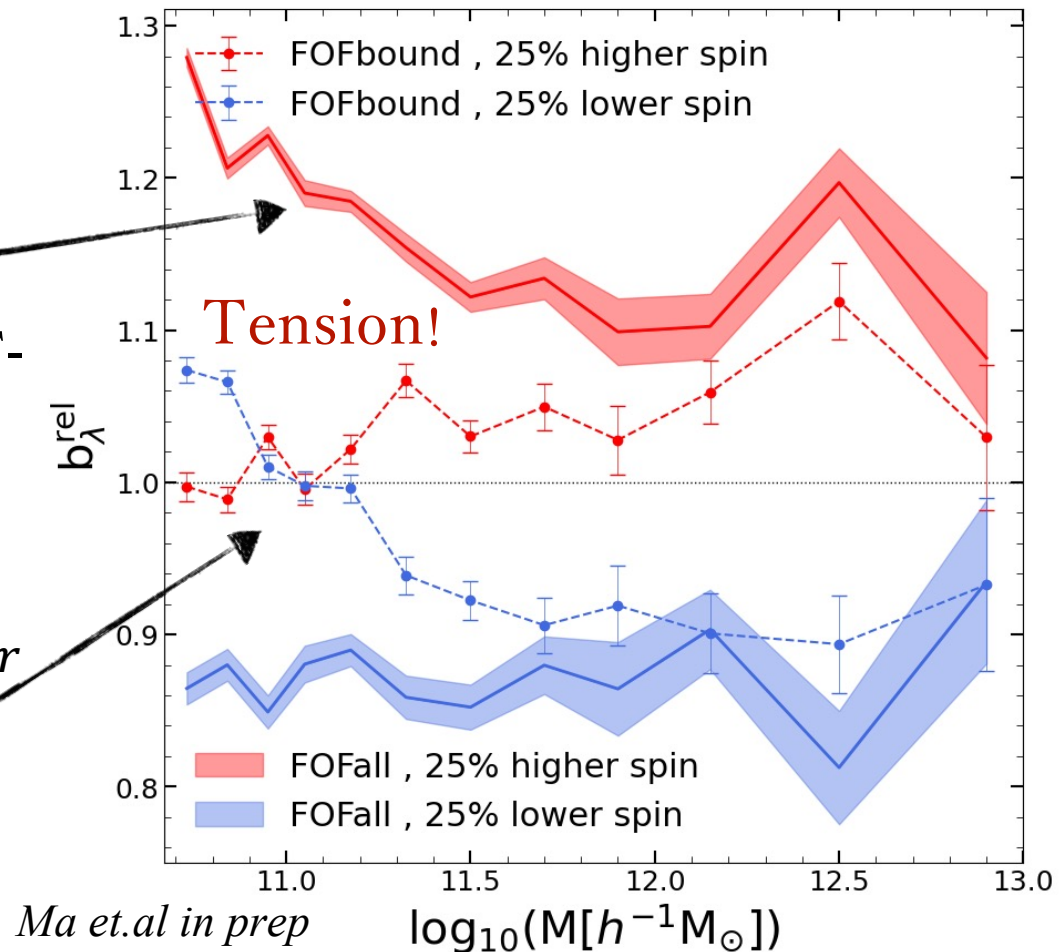
## FOFall:

All particles within  $R_{vir}$  in the FOF-group.

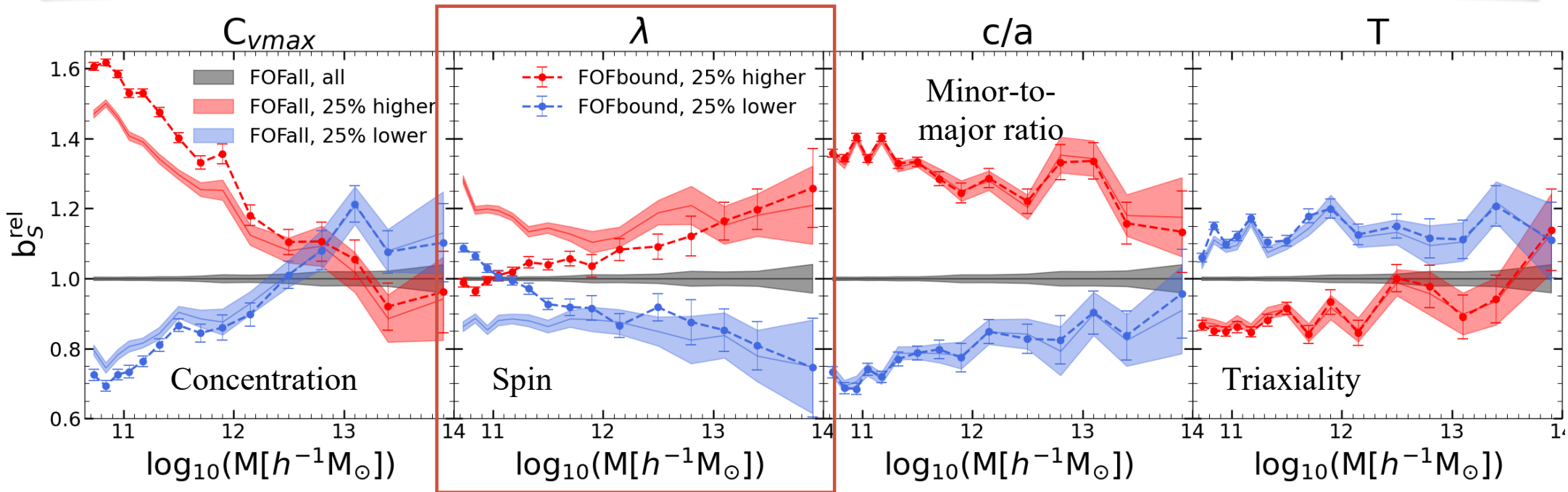
## FOFbound:

Gravitationally bound particles within  $R_{vir}$  in the FOF-group.

Bound: kinematic energy < potential energy



**For different halo properties:** spin bias is strongly dependent on the halo definition, while concentration bias shows a weak relation, and shape bias is not affected by the halo definition.



**FOFall:**

All particles within  $R_{vir}$  in the FOF-group.

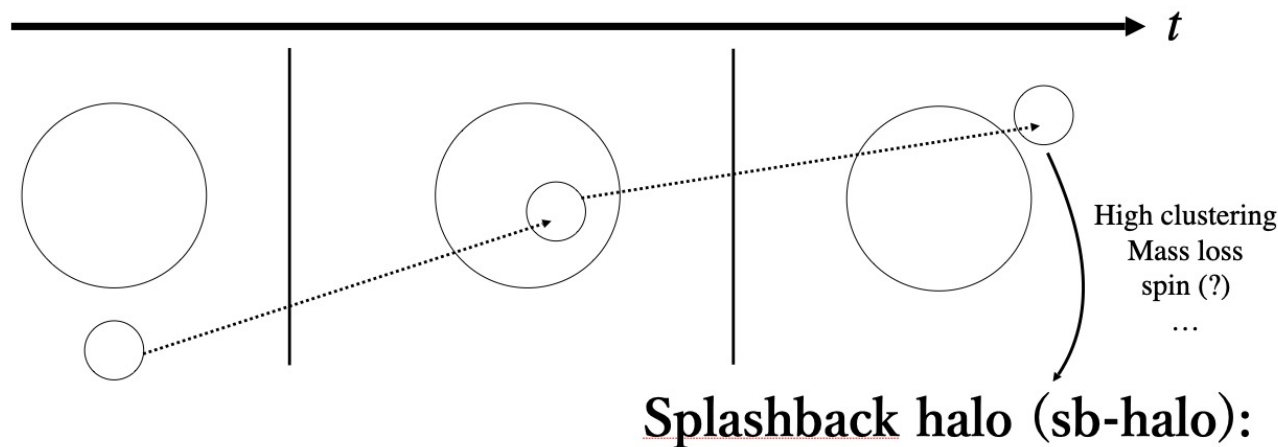
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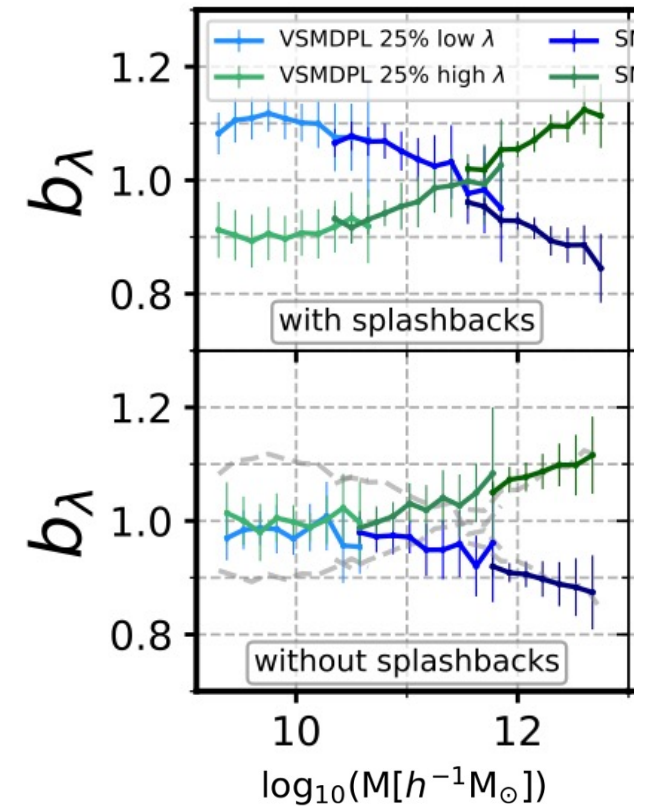
# The physical origin of the low-mass spin bias:

The splashback halo is thought to be **the only reason** for the low-mass spin bias

Calculate the spin by bound particles  $z = 0.0$



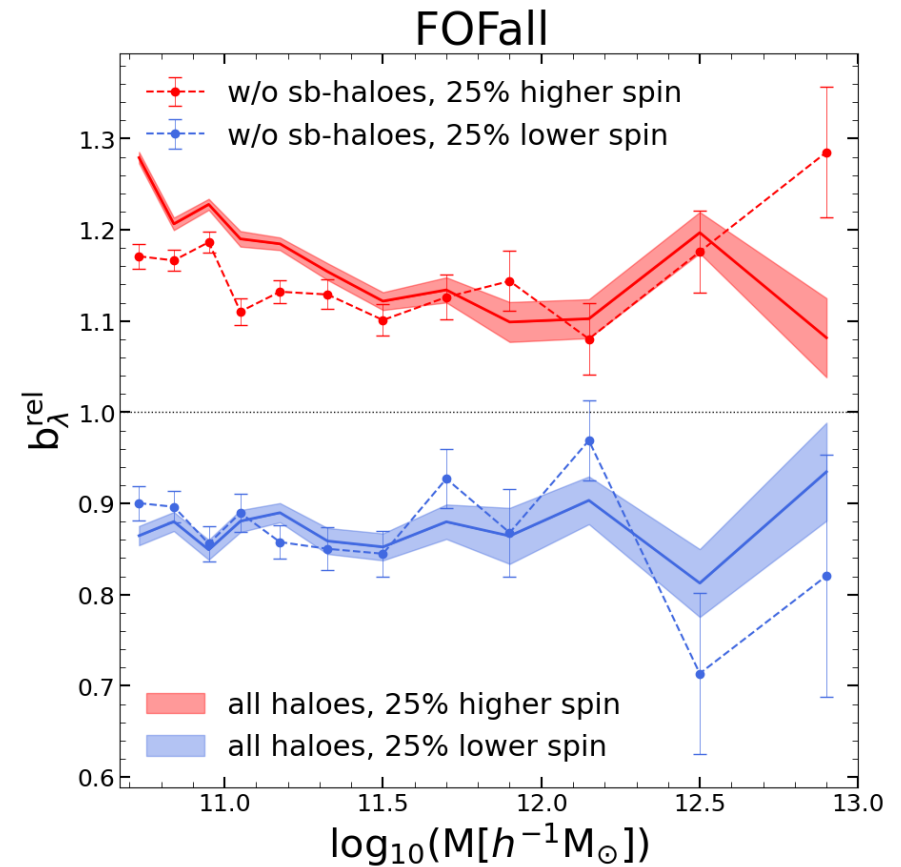
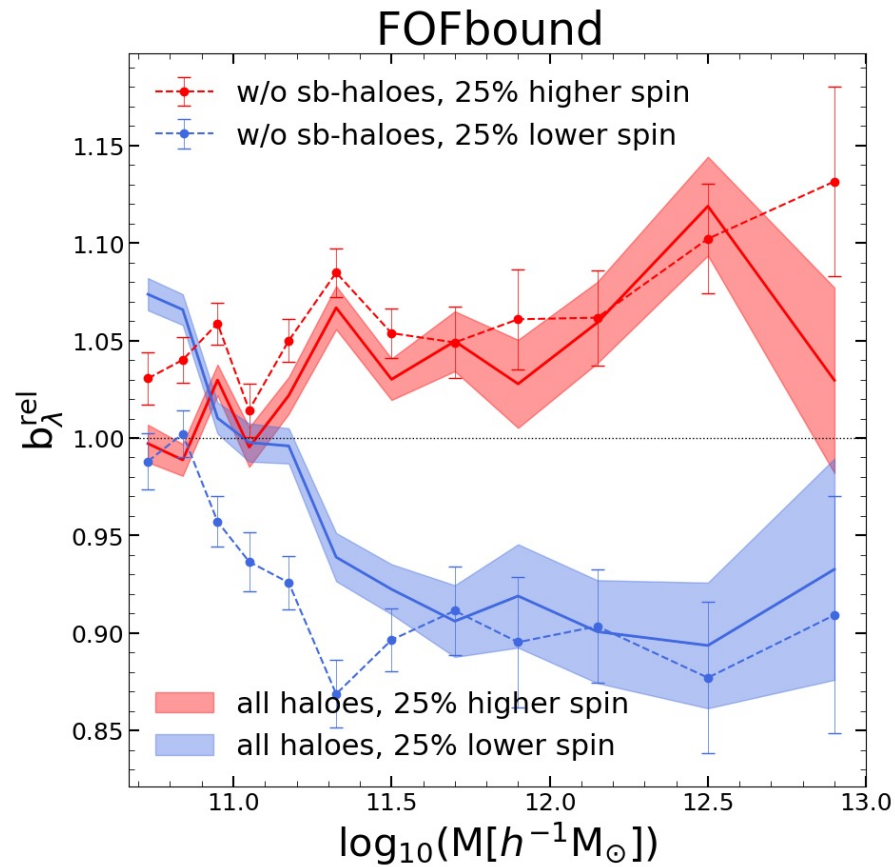
Distinct halo that was subhalo at some previous time, i.e., passed through the virial radius of a larger halo.



*Montero-Dorta et.al 2020; Tucci et.al 2021*

# Spin bias is existed for FOFall at low mass for non-sb haloes:

The splashback halo is thought to be **the only reason** for the low-mass spin bias?

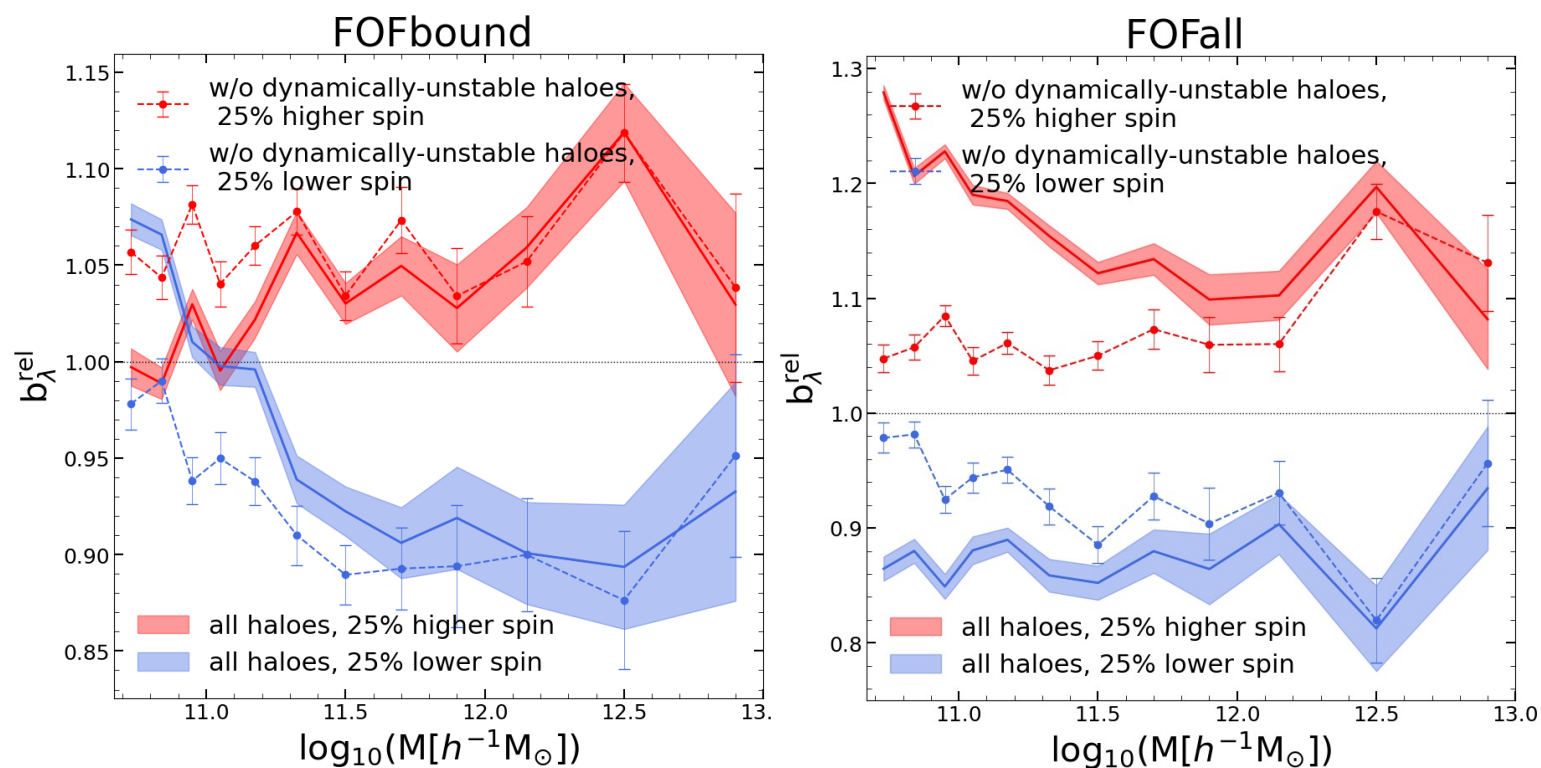




# The physical origin of the spin bias:

Two mechanisms (dynamically unstable/stable halo) of spin bias

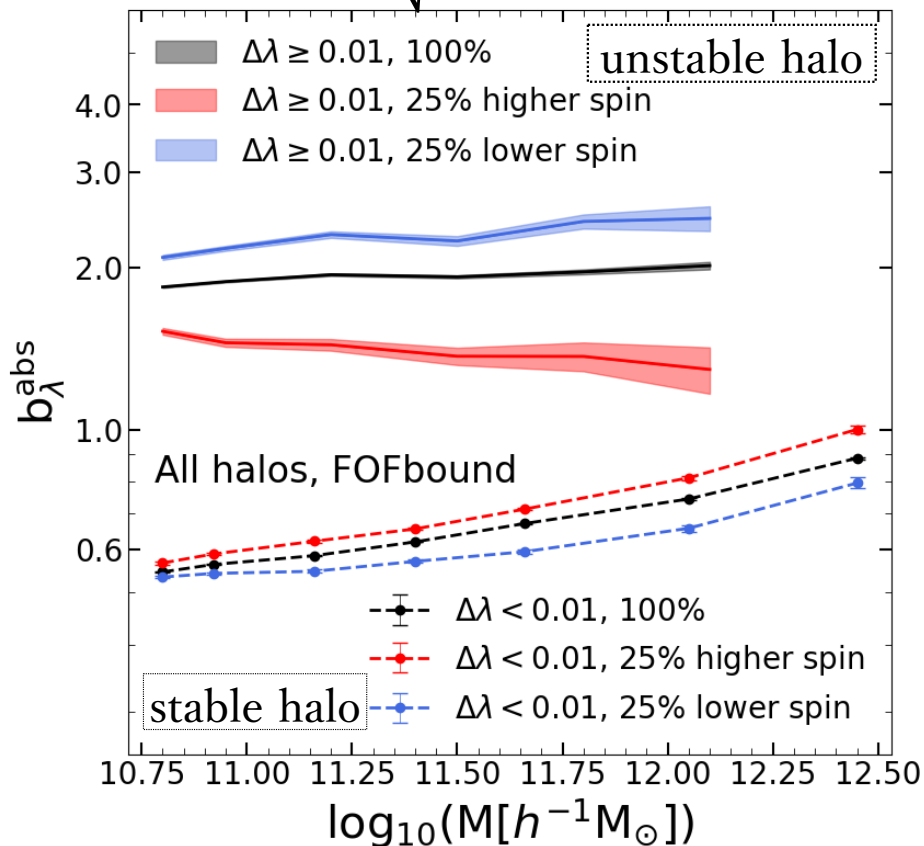
$\lambda_{all}$ : all particles  
 $\lambda_{bound}$ : grav-bound particles  
 $\lambda_{all} - \lambda_{bound} = \Delta\lambda$   
 $\Delta\lambda \geq 0.01$  (unstable)  
 $\Delta\lambda < 0.01$  (stable)



If dynamically unstable halo is removed, the residual spin bias is same under any halo definition.

# Two mechanisms of the spin bias

$$b_{\lambda}^{abs}(r) = \sqrt{\frac{\xi_{hh}(r|M_h, S)}{\xi_{mm}(r)}}$$



$\lambda_{all}$ : all particles

$\lambda_{bound}$ : grav-bound particles

$$\lambda_{all} - \lambda_{bound} = \Delta\lambda$$

$\Delta\lambda \geq 0.01$  (unstable)

$\Delta\lambda < 0.01$  (stable)

Comparison between **unstable halo/stable halo**:

- ◆ The unstable halo has higher clustering than the stable halo

**Unstable halo (5%):**

- ◆ low-rotated halo has higher clustering. (Strong tidal field...)

**Stable halo:**

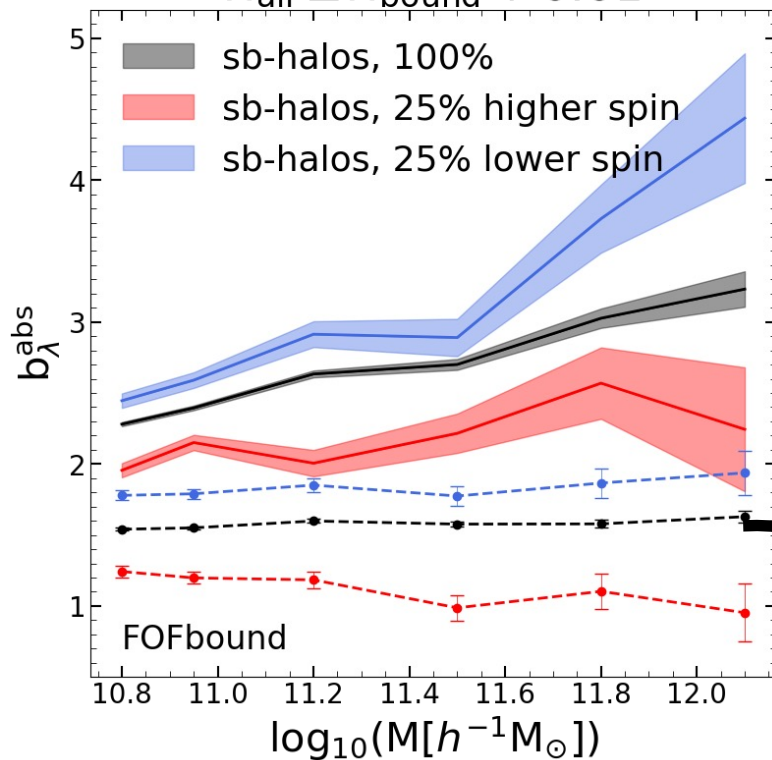
- ◆ high-rotated halo has higher clustering. (Twin halo... Johnson et.al 2019)

# Two mechanisms of the spin bias

Dynamically unstable halo = splashback halo?

Only 40 % of them are splashback haloes.

Dynamically unstable halo  
 $\lambda_{\text{all}} \geq \lambda_{\text{bound}} + 0.01$



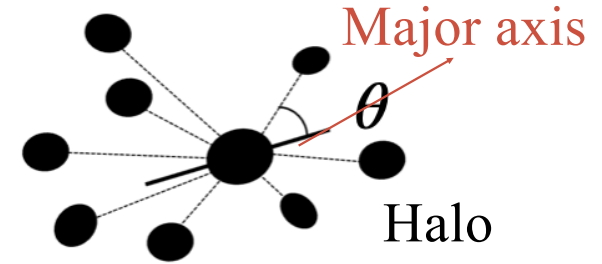
The splashback halo is one of the populations of the unstable halo.

◆ Note the definition of the splashback halo could affect the result.

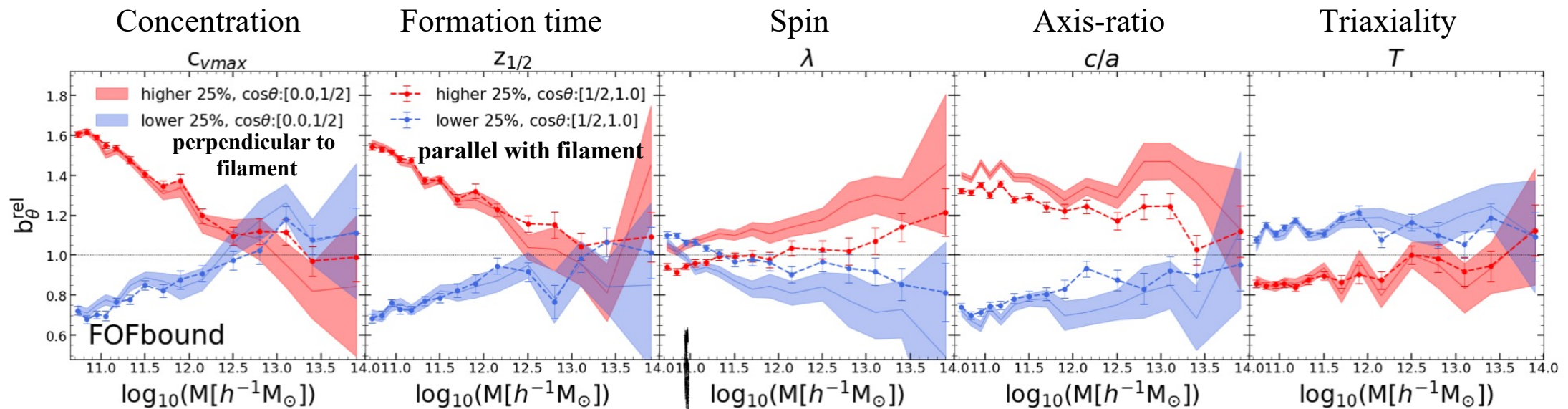
- w/o sb-halos, 100%
- w/o sb-halos, 25% higher spin
- w/o sb-halos, 25% lower spin

# Anisotropic assembly bias:

Strong anisotropy of the assembly bias is shown for the halo spin or the axis-ratio of halos, which indicates that the slow-rotated and elongated halo is more aligned with the filament



$$b_{\theta}^{\text{rel}} = \sqrt{\frac{\xi_{hh}(r|M_{\text{vir}}, S, \theta)}{\xi_{hh}(r|M_{\text{vir}}, \theta)}}$$



*Ma et.al in prep*

The spin bias becomes weak along the major axis (filament)

# Summary

- **Halo definition:**

- (1) The **tension** of the spin bias is caused by the halo definition
- (2) Halo definition largely affects the spin bias, weakly varies the concentration bias, but does not change the shape bias.
- (3) The spin bias is from two kinds of mechanisms: the dynamically stable halo with an intrinsic feature and the dynamically unstable halo with an inverted feature.

- **Spatial anisotropy:**

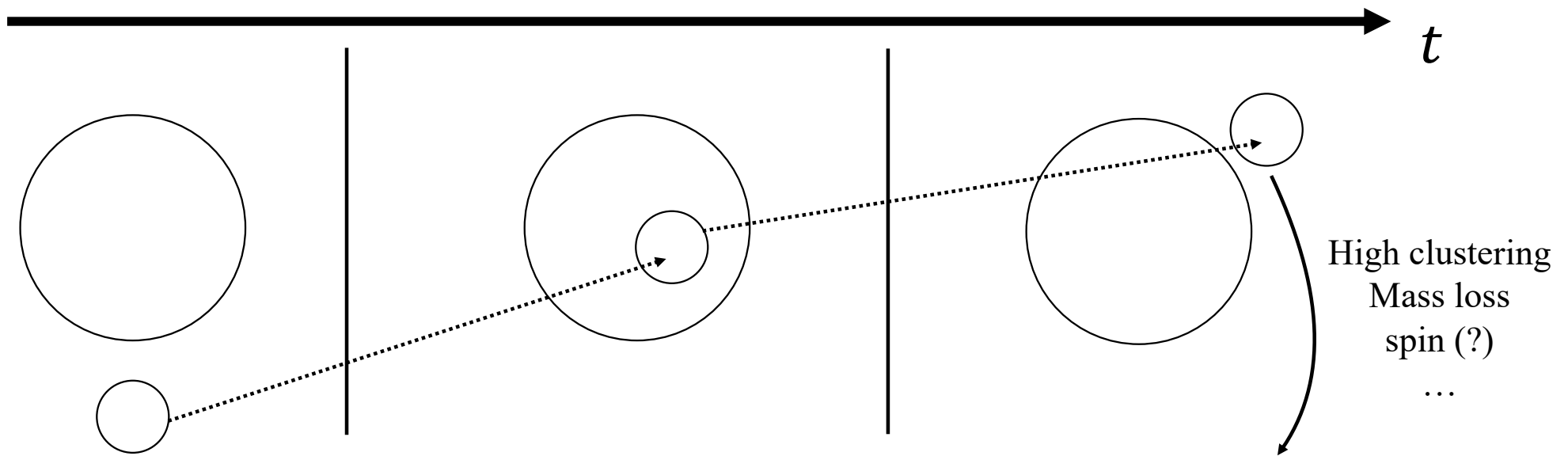
- (1) Strong anisotropy of the assembly bias is shown for the halo spin or the axis ratio of halos, while there is no anisotropy for other properties

# How anisotropic bias reflects the alignment

$$b_{\theta}^{rel} = \sqrt{\frac{\xi_{hh}(r|M_h, S, \theta)}{\xi_{hh}(r|M_h, \theta)}} \quad b_{\theta}^{rel, alignment} = \sqrt{\frac{\xi_{hh}(r|M_h, S, \theta)}{\xi_{hh}(r|M_h, S)}}$$

$$\frac{b(\theta_1, \lambda_2)}{b(\theta_2, \lambda_2)} = \sqrt{\frac{\xi(\theta_1, \lambda_2)/\xi(\theta_1)}{\xi(\theta_2, \lambda_2)/\xi(\theta_2)}} = \sqrt{\frac{\xi(\theta_1, \lambda_2)/\xi(\lambda_2)}{\xi(\theta_2, \lambda_2)/\xi(\lambda_2)}} * \sqrt{\frac{\xi(\theta_2)}{\xi(\theta_1)}} = \frac{b^{alignment}(\theta_1, \lambda_2)}{b^{alignment}(\theta_2, \lambda_2)} * \sqrt{\frac{\xi(\theta_2)}{\xi(\theta_1)}}$$

$$\frac{b^{alignment}(\theta_1, \lambda_2)}{b^{alignment}(\theta_2, \lambda_2)} = \frac{b(\theta_1, \lambda_2)}{b(\theta_2, \lambda_2)} * \sqrt{\frac{\xi(\theta_1)}{\xi(\theta_2)}}$$



## Splashback halo (sb-halo):

Distinct halo that was subhalo at some previous time, i.e., passed through the virial radius of a larger halo.

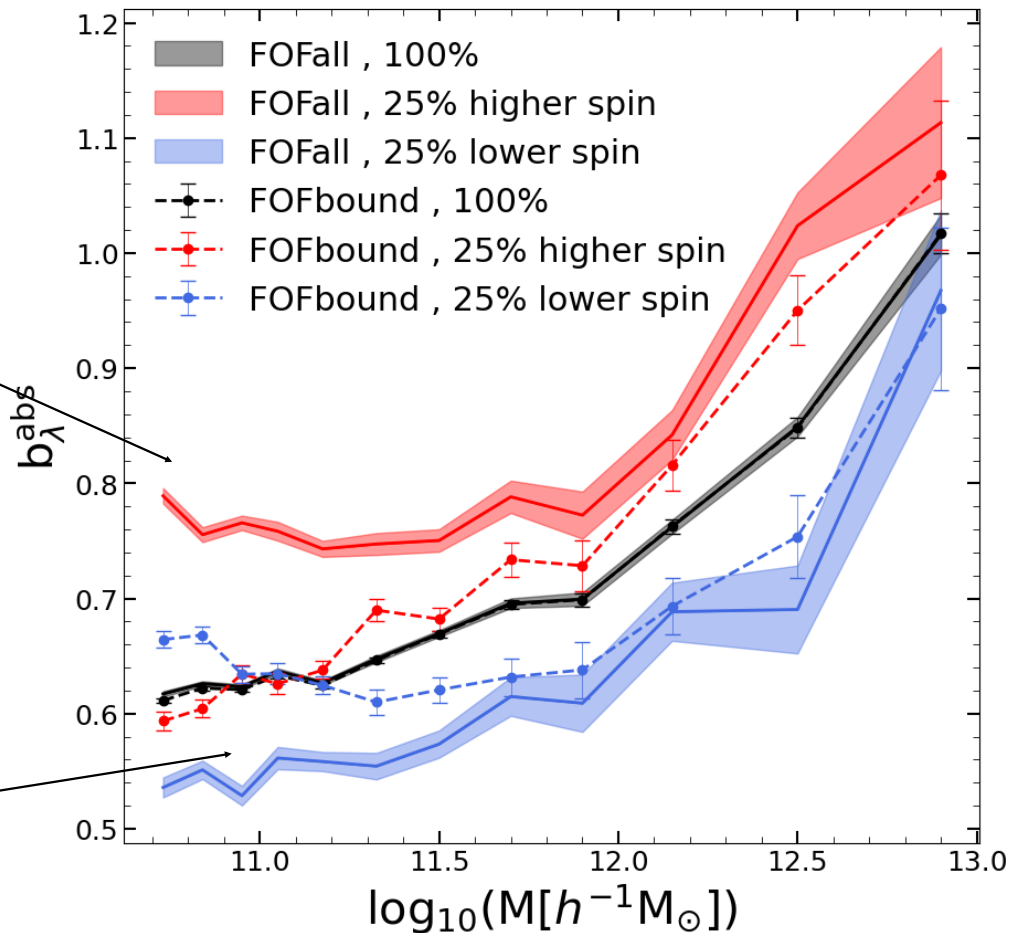
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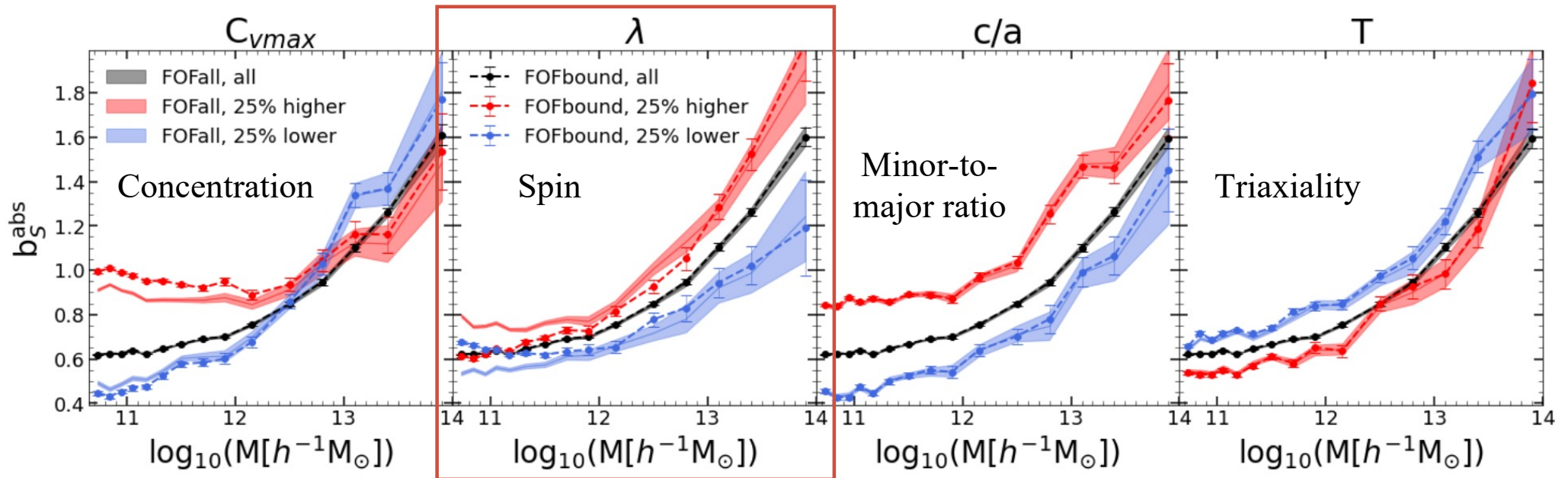
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Gravitationally bound particles within  $R_{vir}$  in the FOF-group.





**For different halo properties:** spin bias is strongly dependent on the halo definition, while concentration bias shows a weak trend and shape bias is not affected by the halo definition.



**FOFall:**

All particles within  $R_{vir}$  in the FOF-group.

**FOFbound:**

Gravitationally bound particles within  $R_{vir}$  in the FOF-group.

# Explanation of anisotropic spin bias

