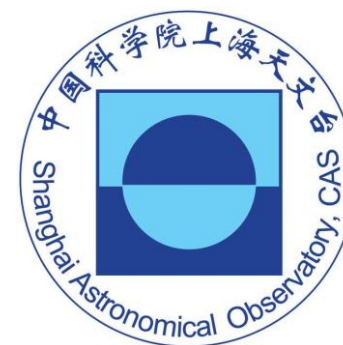


Quantifying the stellar ages of dynamically-separated bulges and disks of CALIFA spiral galaxies

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Collaborators:

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Glenn van de Ven, Shude Mao

2nd Shanghai Assembly on Cosmology and Structure Formation

Background



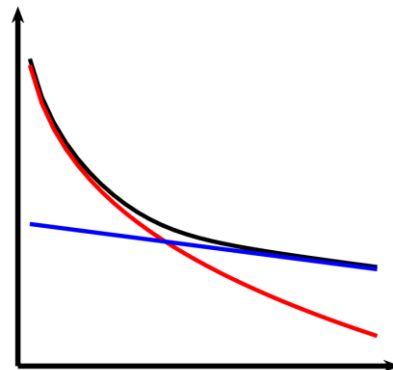
bulge
disk



How these galaxy structures form?

What's the age difference between bulges and disks?

Does this has relation with galaxy fundamental properties?



photometric bulge-disk decomposition
different surface brightness profiles

Background

IFS surveys



spectra across the entire face
of target galaxies



stellar kinematics (V , σ ...)
stellar population (t , Z ...)

Population-orbit superposition method



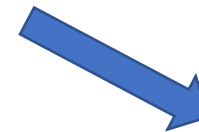
Surveys based on MUSE

high resolution

individual galaxies

(Poci+2019,2021, Zhu+2020,2022a,b, Ding+2023)

Yuchen's talk this morning / Ling's previous talk

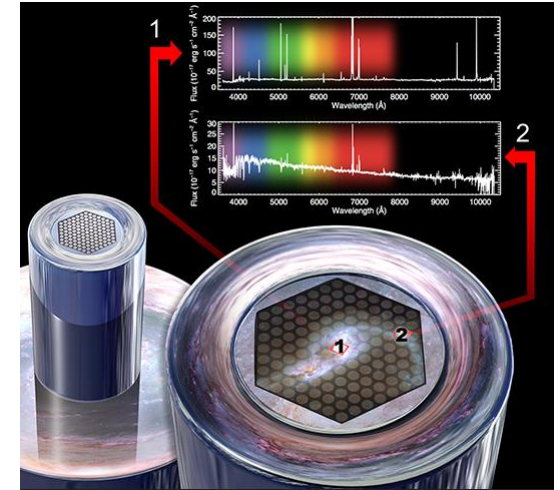


MaNGA, CALIFA, SAMI, ...

medium resolution

large samples

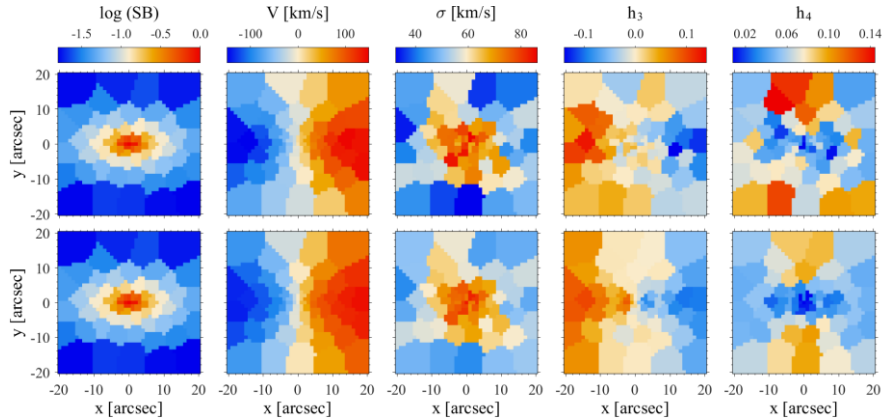
This work



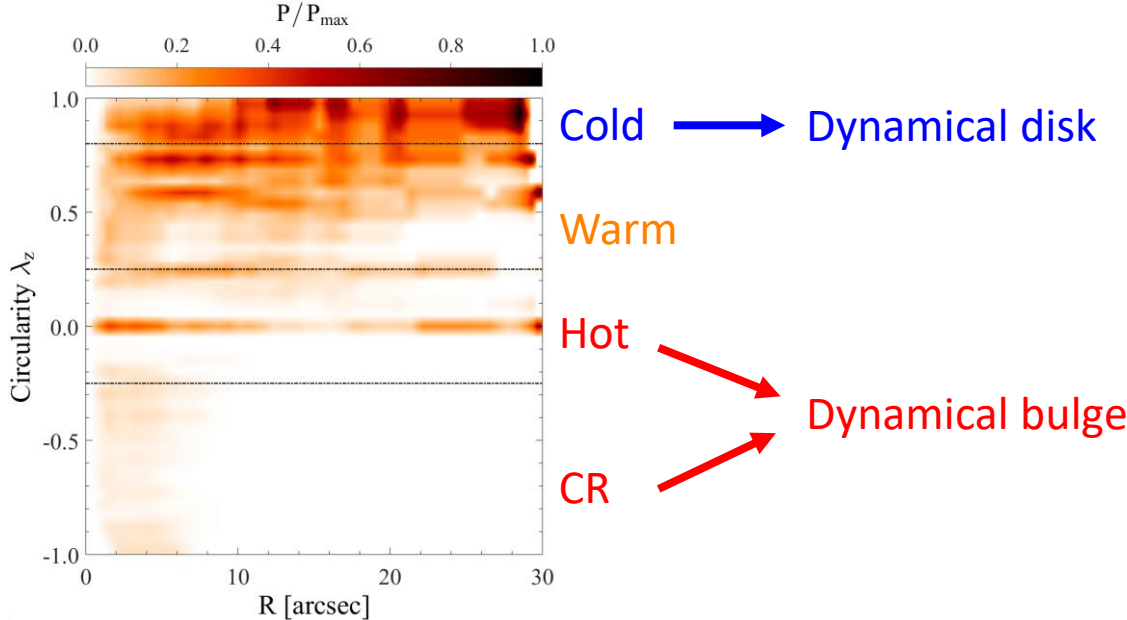
The population-orbit superposition method

① Obtain the stellar orbit distributions by fitting SB and kinematic maps

Observation



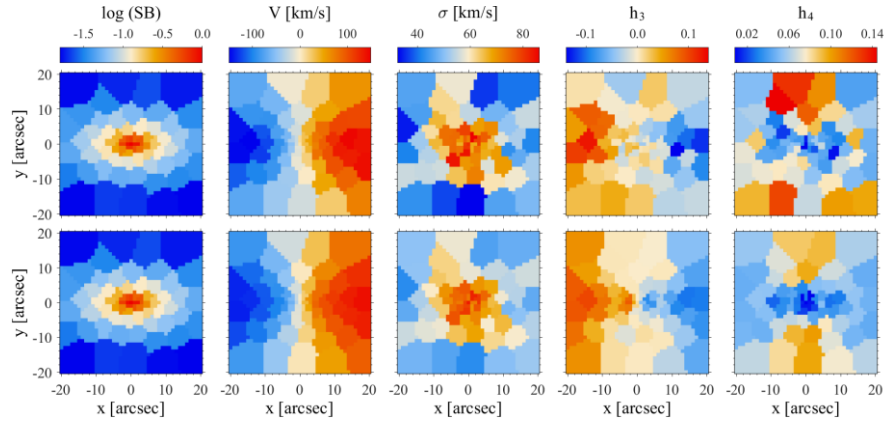
Model



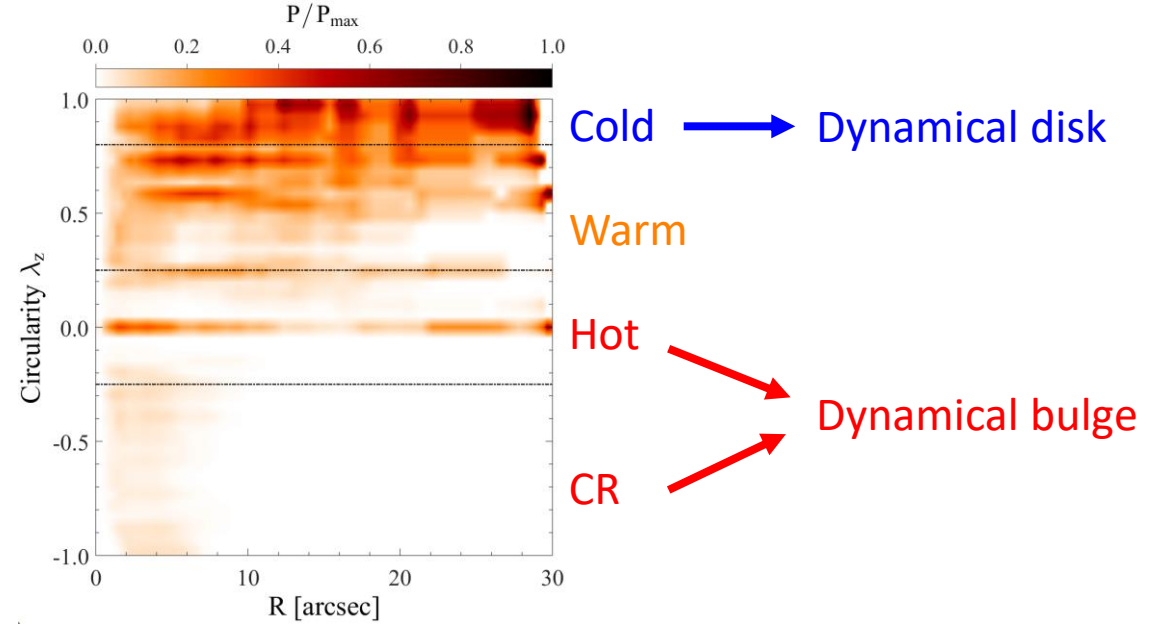
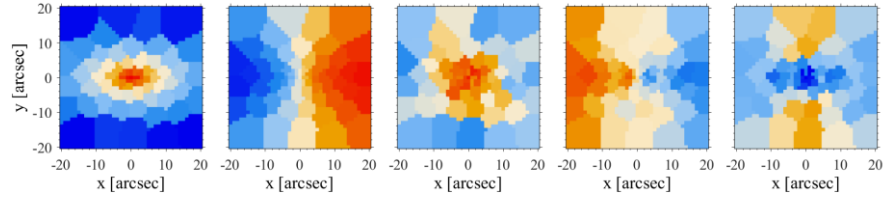
The population-orbit superposition method

- ① Obtain the stellar orbit distributions by fitting SB and kinematic maps

Observation

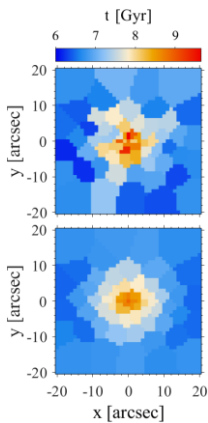


Model

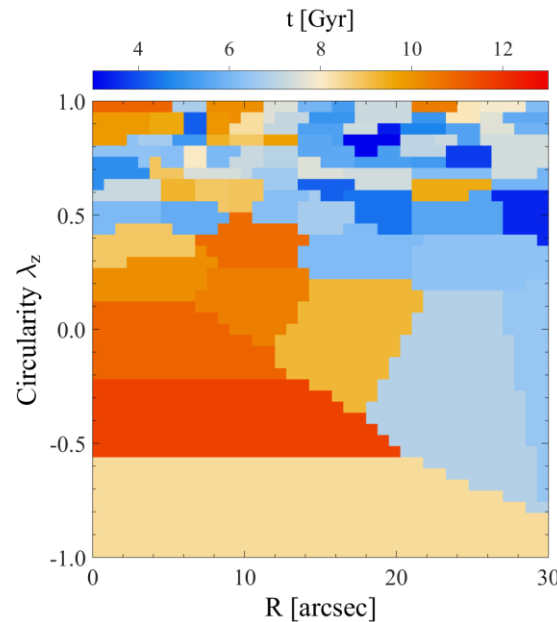
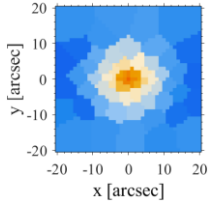


- ② Tagging the stellar orbits with ages by fitting the age map

Observation



Model



The ages of cold/warm/hot+CR components

Disk age
Bulge age

Method validation

3 simulated spiral galaxies from Auriga,
each with 3 different viewing angles

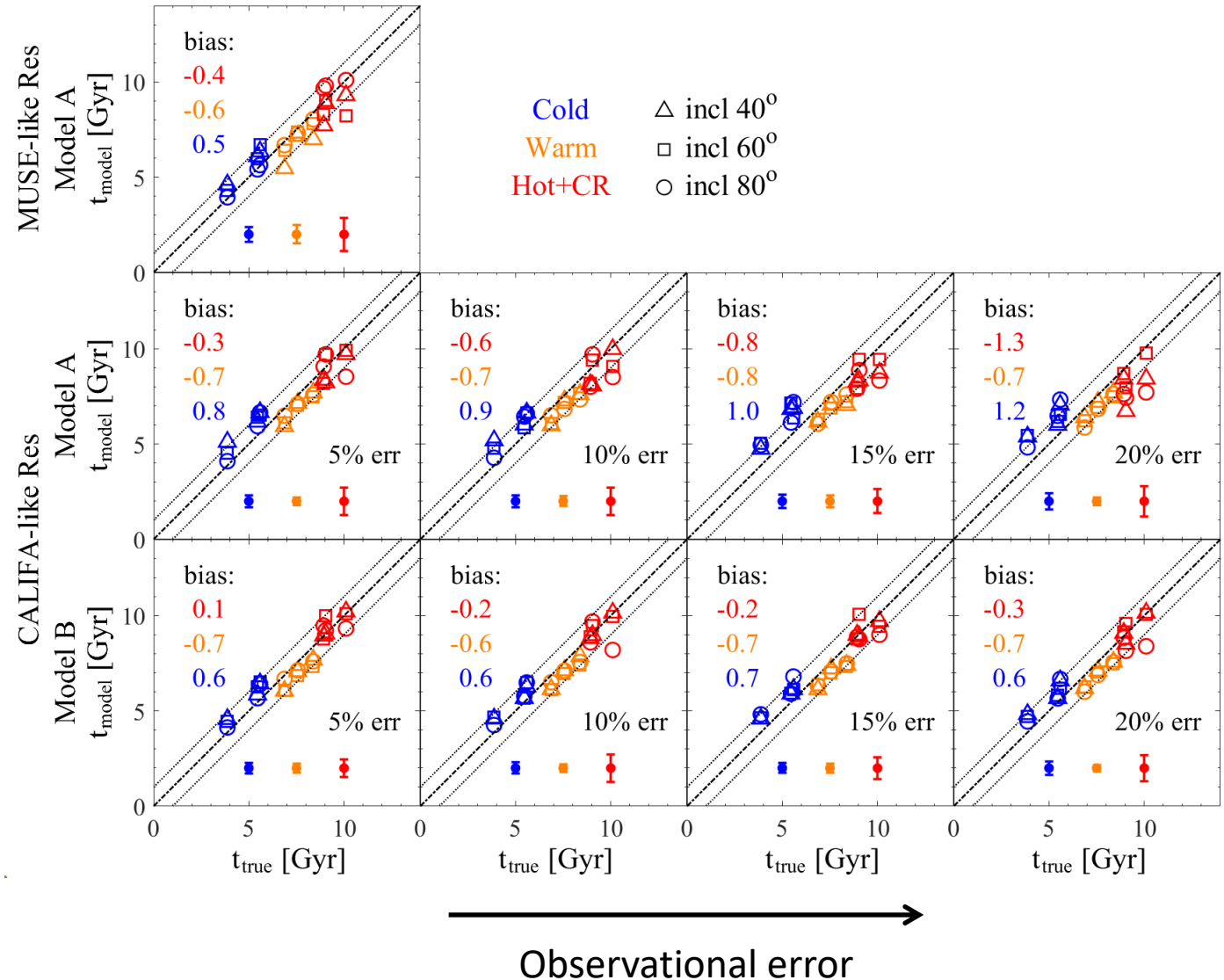
Original model (Model A)
MUSE-like mock data
(Zhu+2020)
Bias < 0.6 Gyr



Original model (Model A)
CALIFA-like mock data
Bias > 1.2 Gyr



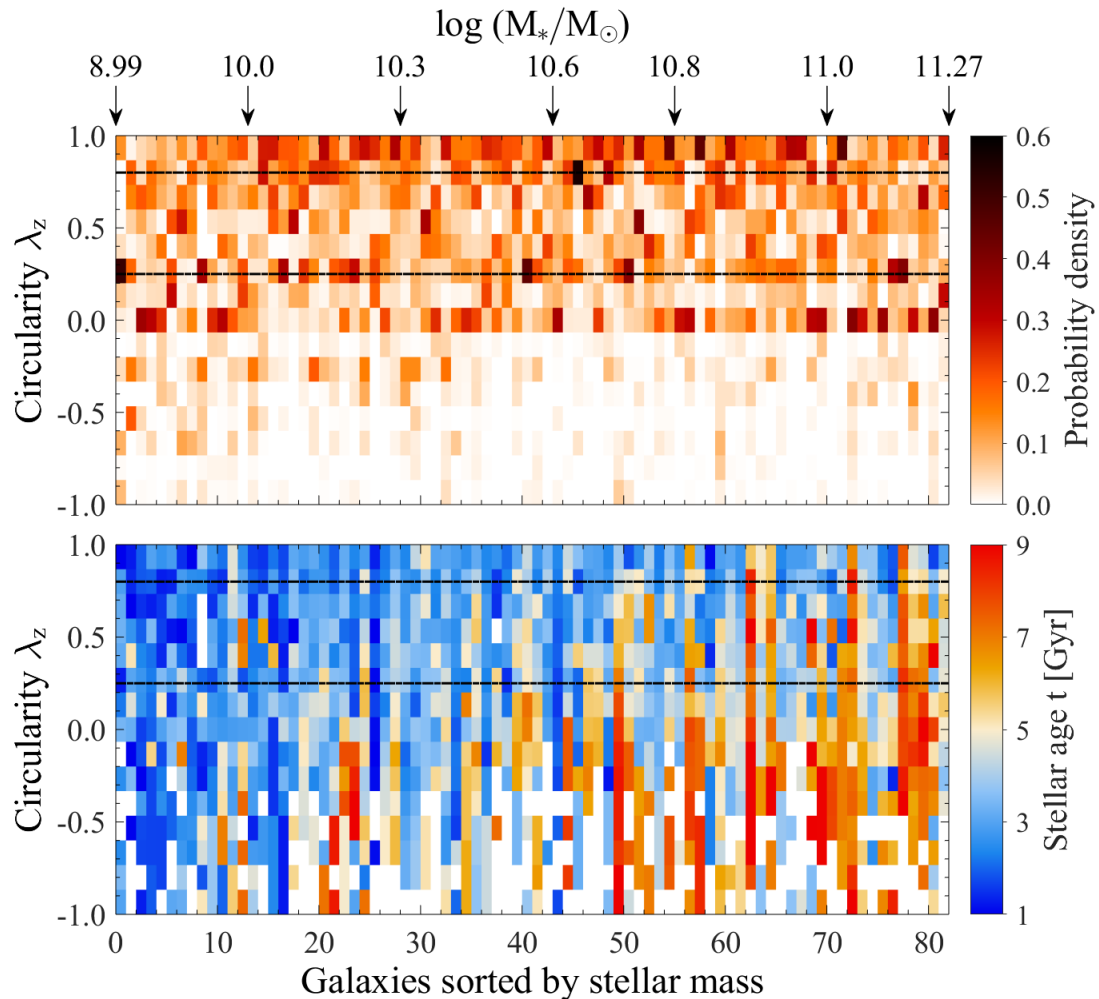
Revised model (Model B)
CALIFA-like mock data
Bias < 0.7 Gyr



CALIFA Results

Apply to 82 CALIFA spiral galaxies

stellar mass: $10^{8.9} \sim 10^{11.3} M_{\text{sun}}$



Stellar orbit distributions $P(\lambda_z)$

Stellar age t



Bulge

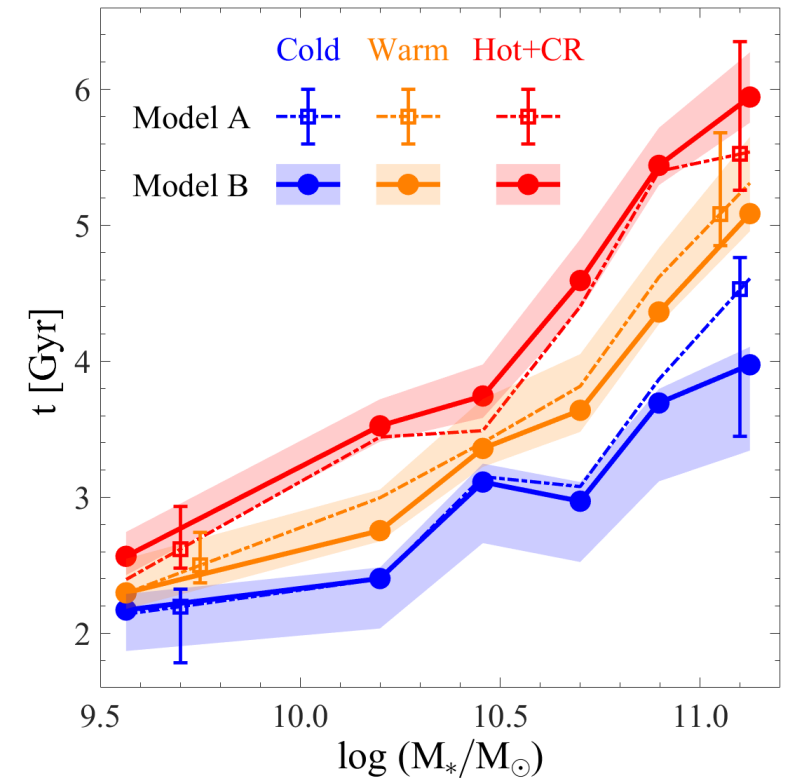


Disk



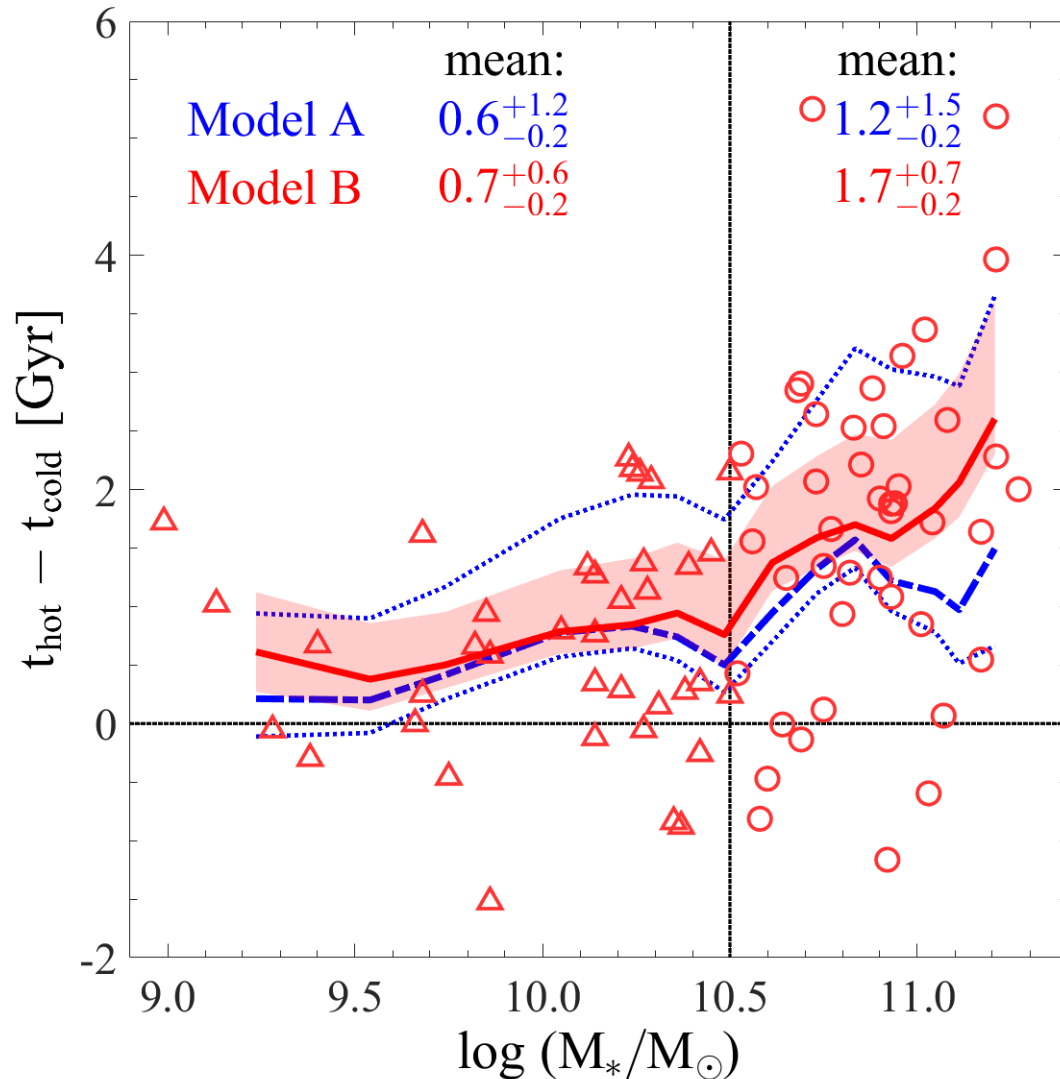
- $t(\text{hot+CR}) > t(\text{warm}) > t(\text{cold})$
- $t(\text{hot+CR})$, $t(\text{warm})$, $t(\text{cold})$ all increase with stellar mass
- The increase of $t(\text{hot+CR})$ is faster than $t(\text{cold})$

Binned ages of different components



CALIFA Results

Age differences between
dynamical bulges and disks



$M_* \sim 10^{10.5} M_{\text{sun}}$:

the increase of $t(\text{hot+CR}) - t(\text{cold})$ become rapid

Lower mass spirals ($M_* \leq 10^{10.5} M_{\text{sun}}$):

young bulge + young disk

form at intermediate
redshift

assembled right after the
formation of bulges
experienced longer period of
star formation

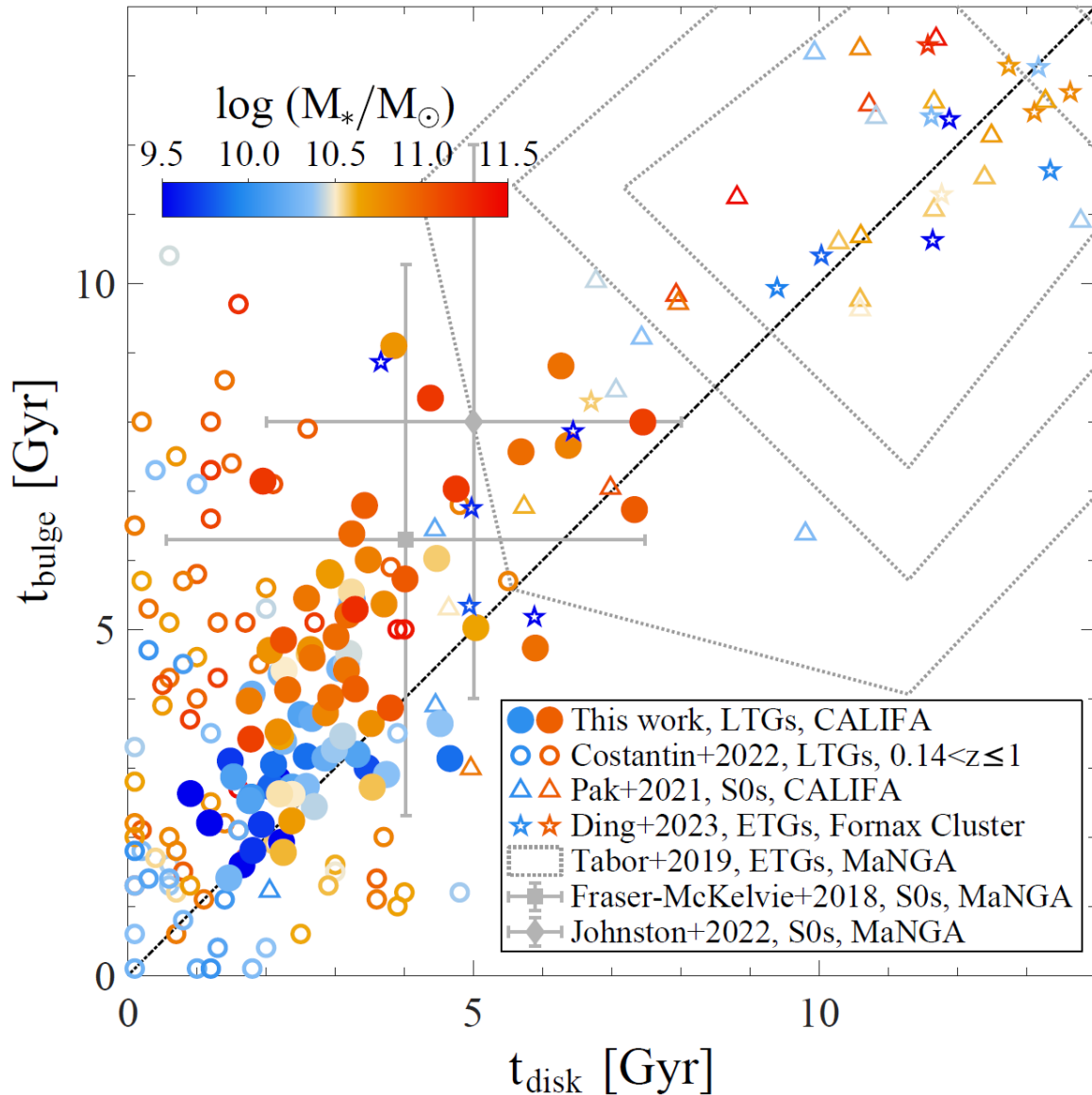
Higher mass spirals ($M_* > 10^{10.5} M_{\text{sun}}$):

old bulge + young disk / old disk

form at high
redshift

different star formation histories
various merger histories

Bulge versus disk stellar ages



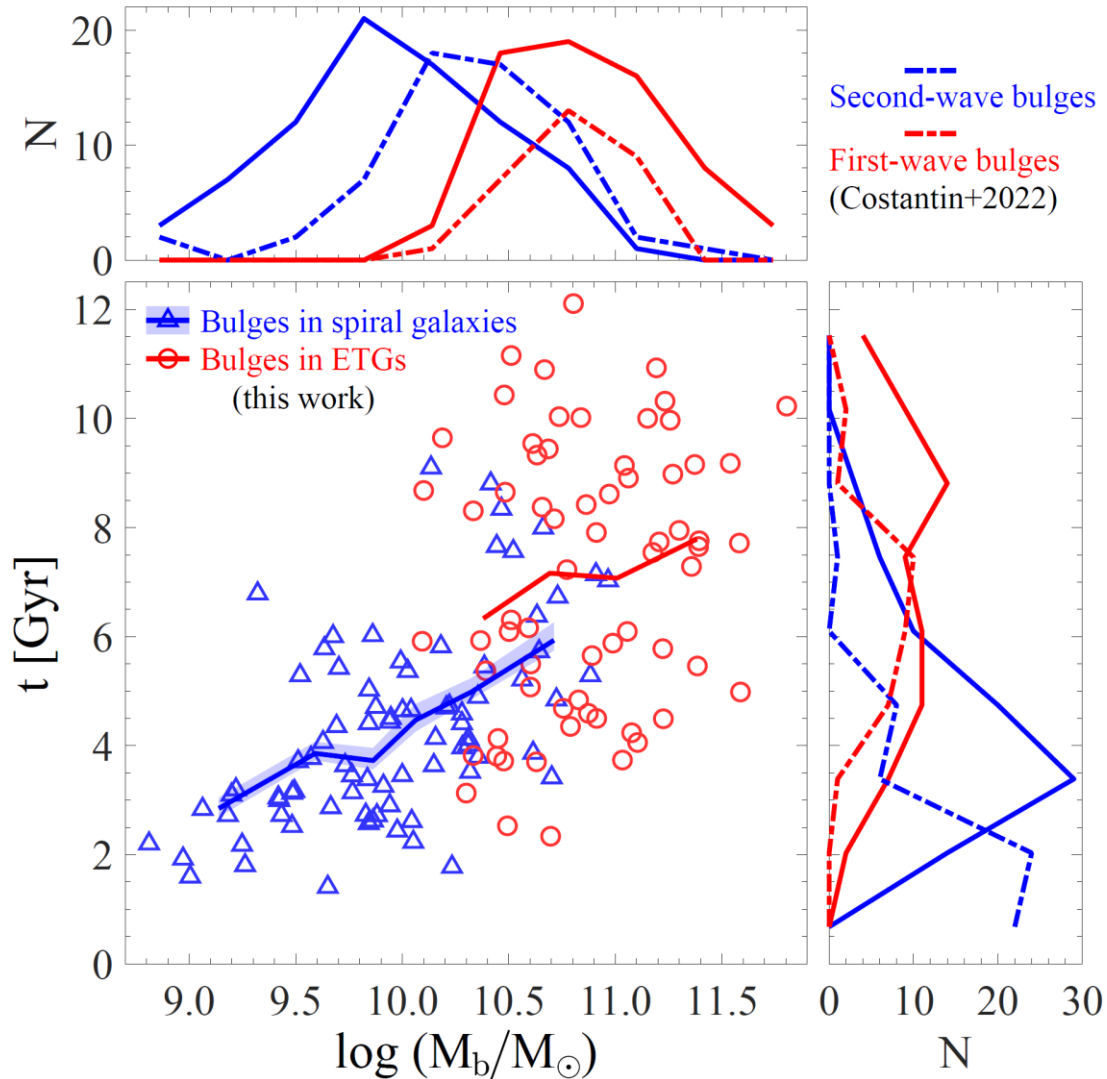
Spiral galaxies in CALIFA/higher redshift:
most galaxies have bulges older than disks
 $t(\text{bulge}) - t(\text{disk})$ increase with stellar mass

S0 galaxies in MaNGA/CALIFA:
bulges are slightly or 2~3 Gyr older than disks

Early-type galaxies in MaNGA:
similar bulge ages and disk ages

Early-type galaxies in the Fornax cluster:
ancient infallers: similar bulge ages and disk ages
recent infallers: bulges are ~ 2 Gyr older than disks
(Yuchen's work)

Bulge stellar age versus bulge mass



This work ($z \approx 0$):

82 spirals in CALIFA

67 ETGs in CALIFA

Costantin+2022 ($0.14 < z \leq 1$):

spirals with first-wave bulges

spirals with second-wave bulges

similar origins

similar origins



The bulges display a bimodal distribution of ages

Summary

- We validate the population-orbit superposition method with CALIFA-like mock data, thus can quantify the luminosity fractions and stellar age of different dynamical structures.
- We apply the method to 82 CALIFA spiral galaxies, and find:
 - (1) Both the dynamical disks and bulges are older in more massive galaxies.
 - (2) The bulges are older than disks in $\sim 80\%$ CALIFA spiral galaxies, with their age difference become larger in more massive galaxies.
 - (3) Combining 82 spirals with 67 ETGs, the bulges in CALIFA galaxies show bimodality in both stellar age and mass distributions, which is consistent with the two-wave bulge formation (Costantin+2022).

Accepted by A&A

<https://arxiv.org/abs/2310.07201>

Thank you!

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